

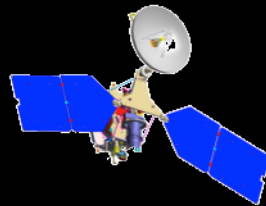


Jet Propulsion Laboratory  
California Institute of Technology

# The Power of the Mars Program

University of Michigan

April 5, 2017



**Leslie K. Tamppari**  
Deputy Project Scientist,  
Mars Reconnaissance  
Orbiter

Jet Propulsion Laboratory, California Institute of Technology

# Outline

- Intro slide and my background
- Overview of Mars Exploration Program (5)
- Interaction between the projects – Intro slide
  - Science, Landing sites section, Aerobraking, Landing support, Relay
- MRO is key player
  - Instrument overview
  - How these data are used in
    - Science
    - Landing site selection and safety
    - Aerobraking and landing
  - Recent results
    - ExoMars EDM
    - InSight
    - ExoMars
    - 2020 (NASA and ESA)
    - Red Dragon
- NASA 2020 – three final sites
  - Show them
  - Science overview of the mission
  - Instrumentation
- Summary

# 4 MARS SCIENCE GOALS

1

WATER

## LIFE

Determine if life ever arose on Mars

2

HABITABLE ZONES

## CLIMATE

Understand Martian climate processes and history

3

SIGNS OF LIFE

## GEOLOGY

Determine how the surface and interior of Mars evolved

4

## HUMANS

Prepare for human exploration



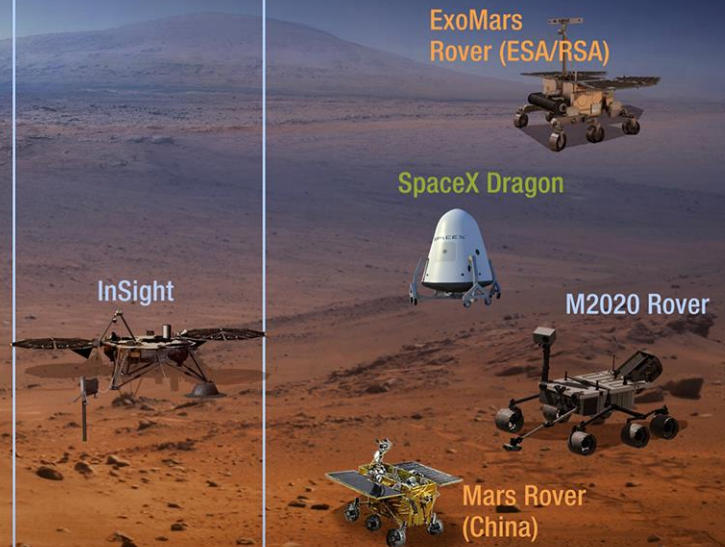
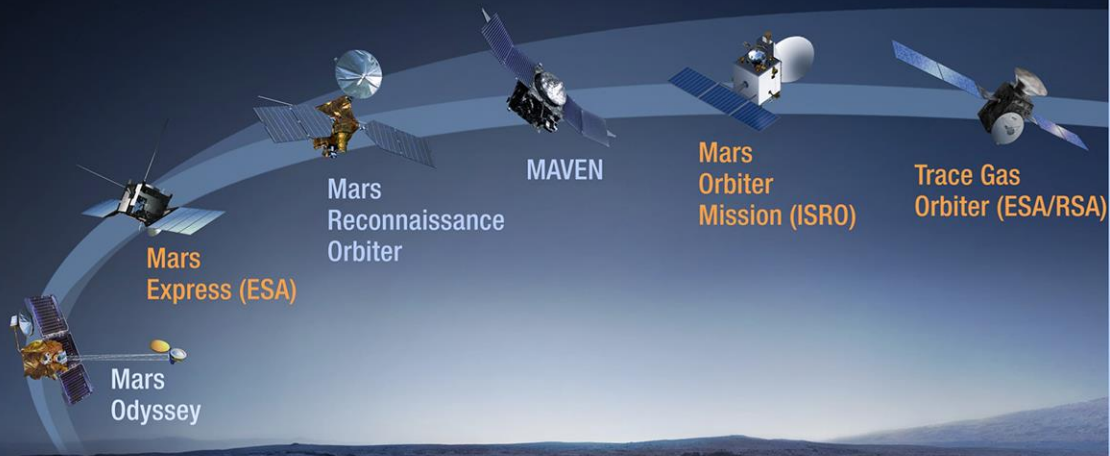
EVOLVING THEMES



Operational 2001–2017

2018

2020 and Beyond



*Follow the Water*

*Explore Habitability*

*Seek Signs of Life*

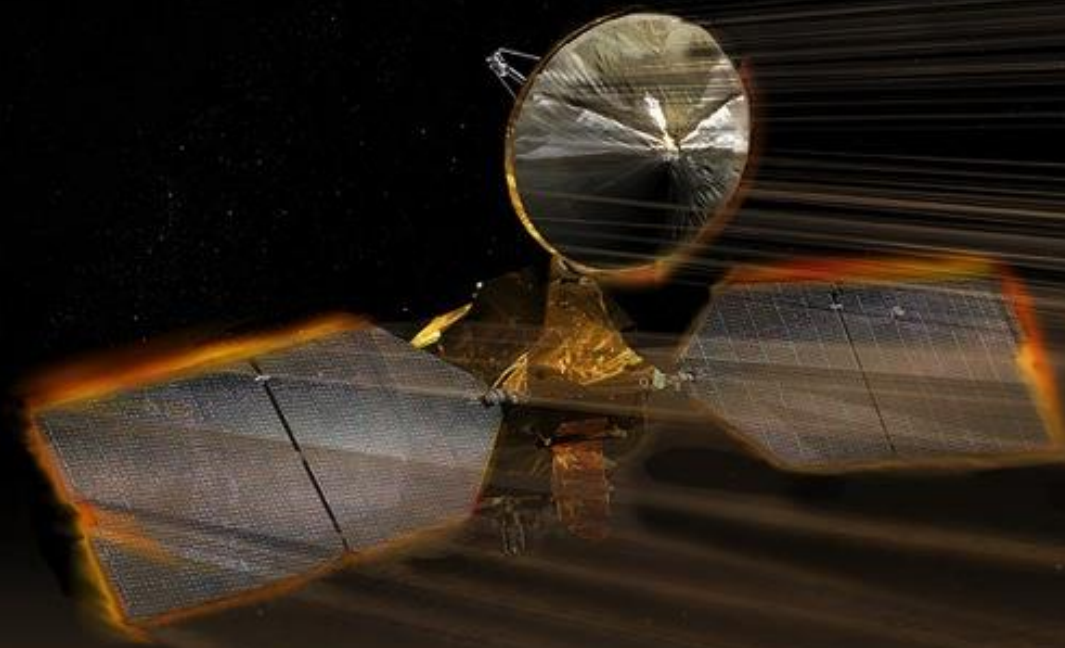
*Prepare for Future Human Explorers*



# Interaction between spacecraft

- Science, Landing sites seection, Areobraking, Landing support, Relay

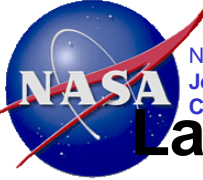
JPL 2006



*Artist's Concept*

*Credit: LMSSC / JPL / NASA*





National Aeronautics and Space Administration  
Jet Propulsion Laboratory  
California Institute of Technology

Mars Reconnaissance Orbiter

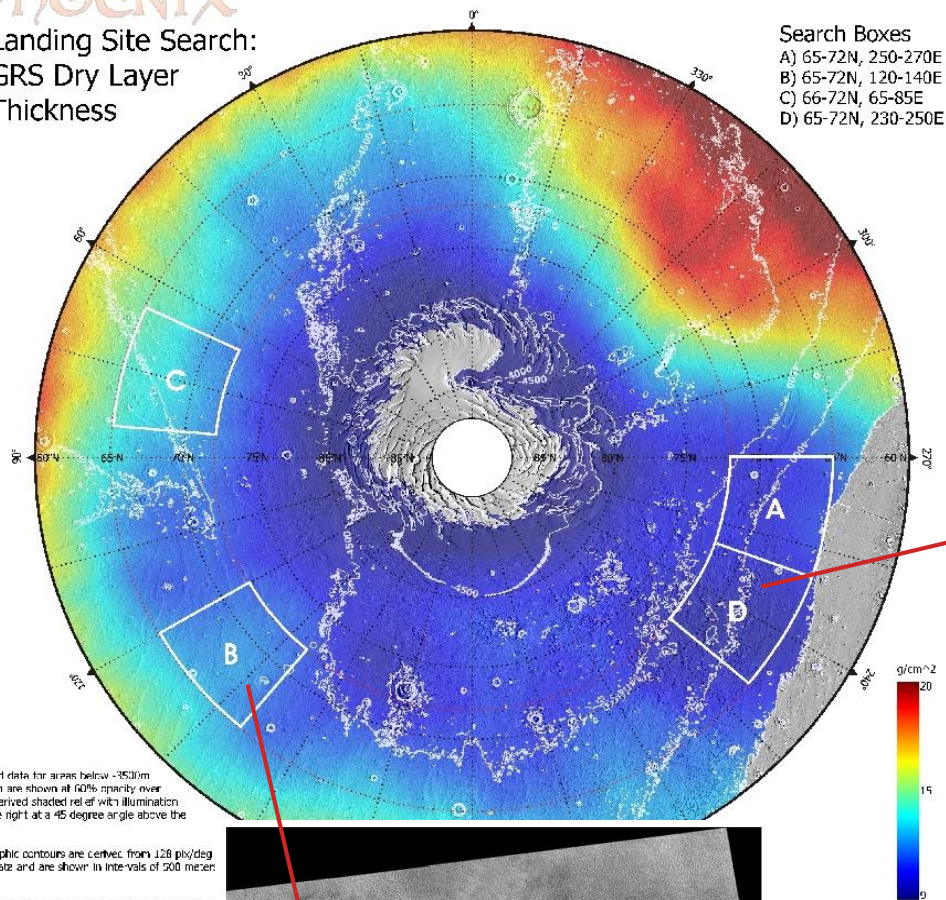
# Support of other missions:



## Landing site selection, characterization, certification

PHOENIX

Landing Site Search:  
GRS Dry Layer  
Thickness

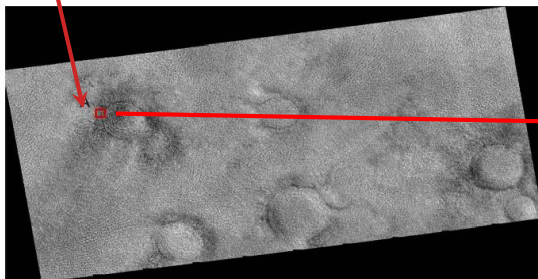
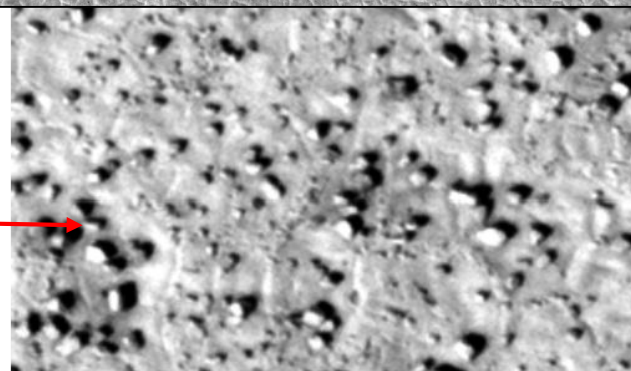


Search Boxes  
A) 65-72N, 250-270E  
B) 65-72N, 120-140E  
C) 66-72N, 65-85E  
D) 65-72N, 230-250E

HiRISE/UA/JPL/NASA

Rock  
Abundance  
CFA  
< 5%

100 m



February 23, 201

xstart: 5619  
Ystart: 6078



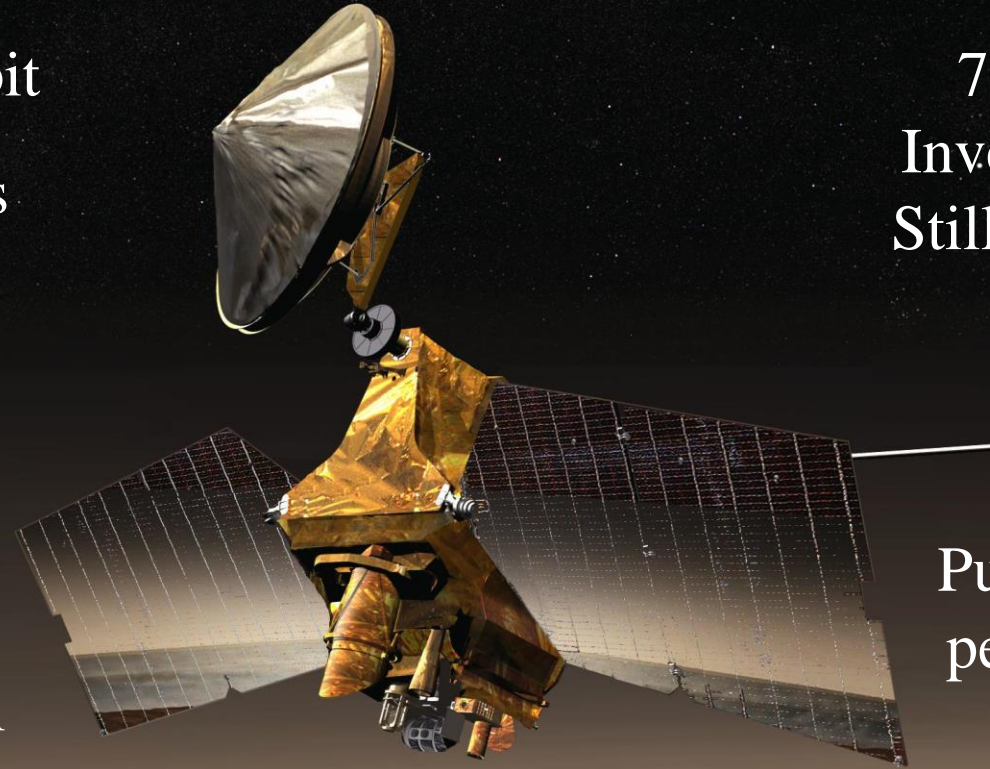
# *Rated “Excellent” in 2016 Planetary Mission Senior Review*

11 Years in Orbit

~50,000 orbits

300 Tb of  
Science Data  
Returned

~200 kg of  
Usable Fuel still  
in the Tank



7 Science  
Investigations  
Still Returning  
Data

~1000  
Publications in  
peer-reviewed  
Journals

## **Mars Reconnaissance Orbiter**

*@Copyright 2016 California Institute of Technology*

*Government sponsorship acknowledged.*

# MRO Science Investigations

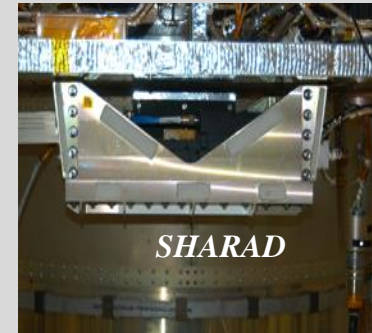


## HiRISE

~49,000 images  
(~5000 stereo)  
~2.8% of Mars  
RSL, Gullies,  
Dunes, Polar Caps

## SHARAD

~21,000 Observing Strips  
Buried CO<sub>2</sub> Ice  
Polar Cap Internal  
Structure  
Mid-latitude Ice



## CRISM

~85% msp IR  
~39% hsp IR  
~76% hsp VNIR  
Limb Scans  
Ancient Aqueous Minerals  
ATO's (6-12 m/pixel)

## MCS

~150 M Soundings  
~94% of 5.4 MYrs  
Vert. Dust Profiles  
Dust Storm Patterns  
Tidal Structure  
CO<sub>2</sub> snow and frost



## CTX

~92,000 images ~99% of  
Mars 20% in dual  
coverage  
Stratigraphy  
New Impacts

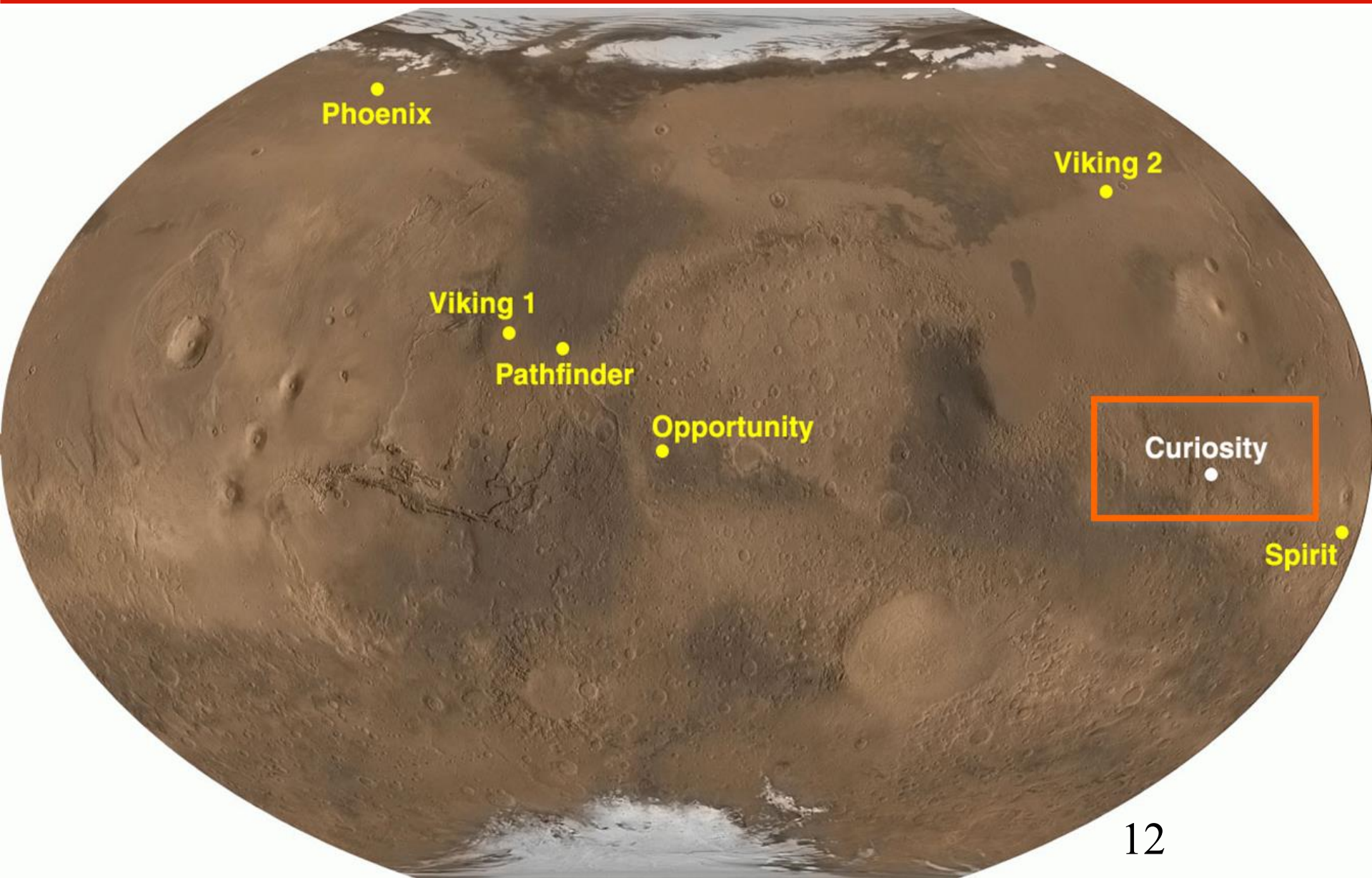
## MARCI

~47,000 images  
5.4 Myrs  
~3600 Daily Global Maps  
Dust & Ice Clouds  
Dust Storm Tracks



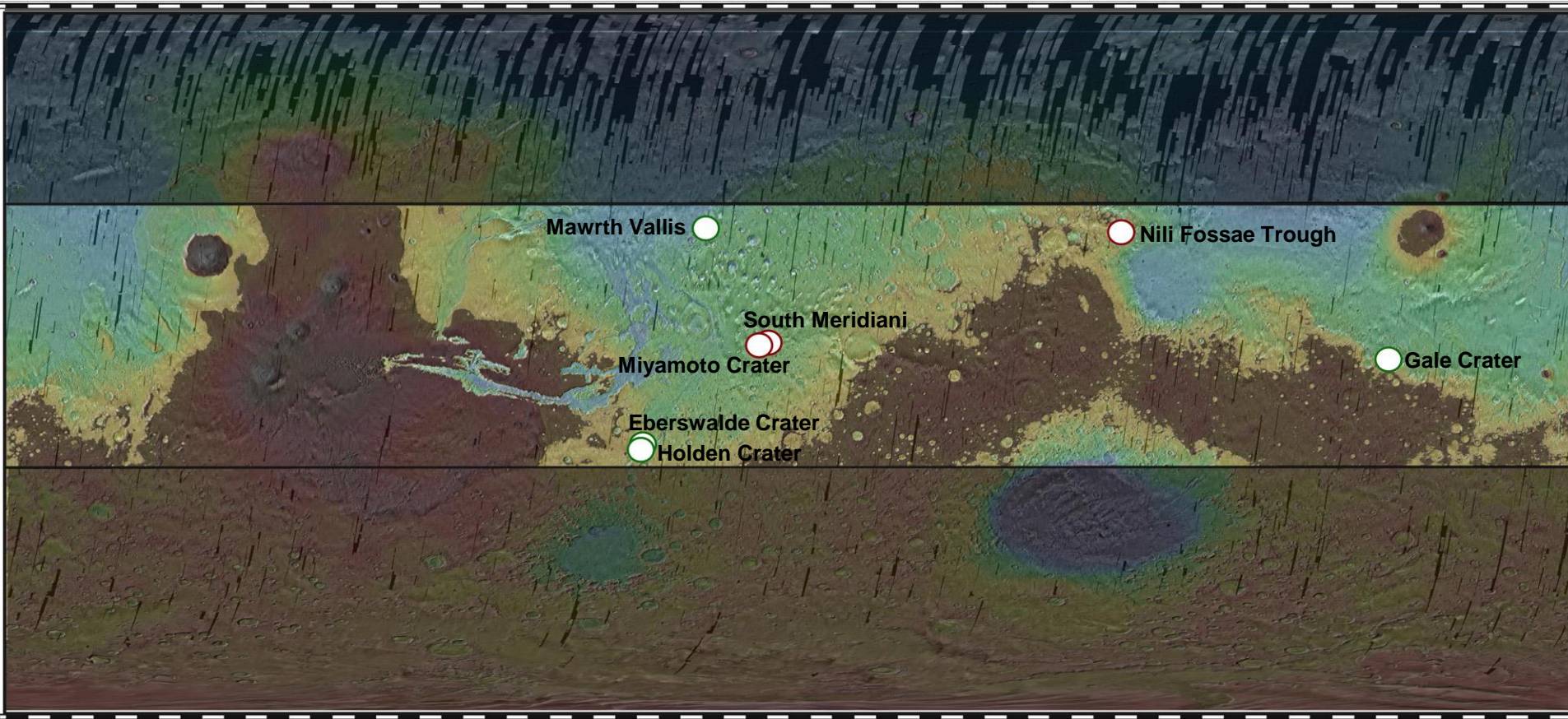


# Mars Landing Sites





## Seven Downselected MSL Landing Sites:



## Seven Sites Receiving Highest Science Ranking:

Shaded areas poleward of  $30^\circ$  , elevations  $>1$  km

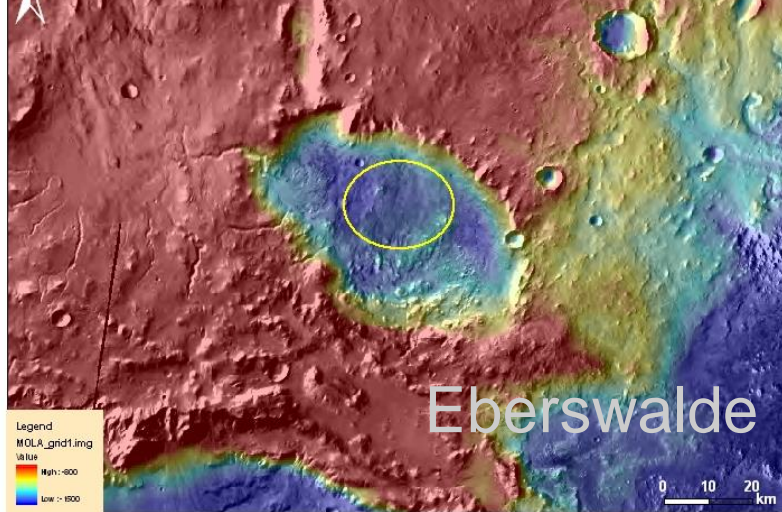
Green outlines denote final four sites based on science, engineering



# Final Four MSL Landing Ellipses

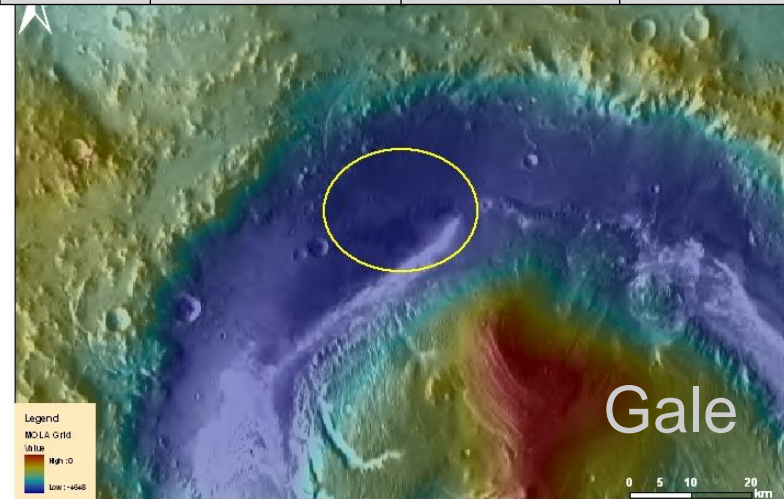
Eberswalde Crater: MOLA

Eberswalde	-23.8953°S	326.7426°E	-1435 m
------------	------------	------------	---------



Gale Crater: MOLA

Gale	-4.4868°S	137.4239°E	-4449 m
------	-----------	------------	---------



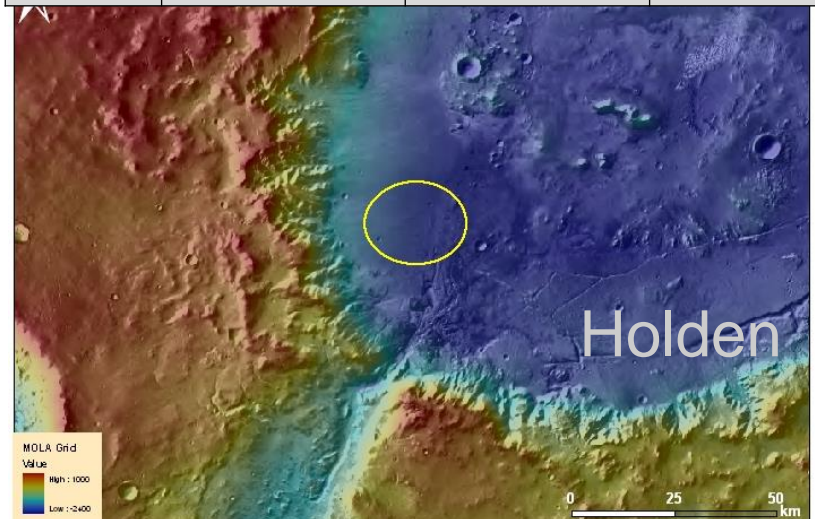
Mawrth Vallis Site 2: MOLA

Mawrth	23.9883°N	341.0399°E	-2231 m
--------	-----------	------------	---------



Holden Crater: MOLA

Holden	-26.4007°S	325.1615°E	-2088 m
--------	------------	------------	---------

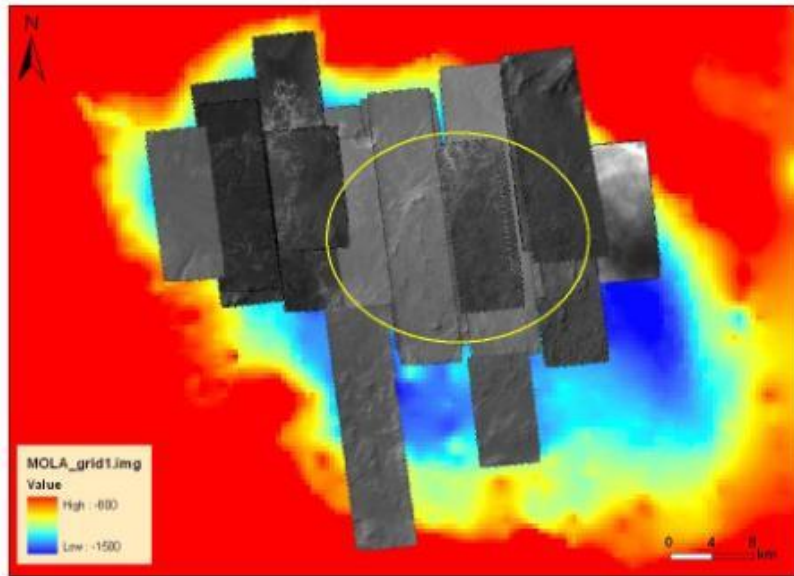


25 km by 20 km Ellipses E-W for 2011

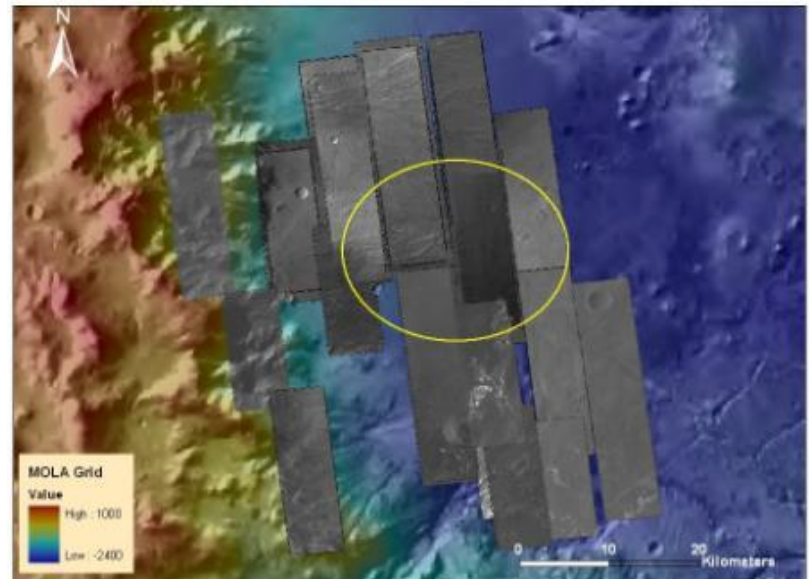


# HiRISE Image Coverage in Ellipse

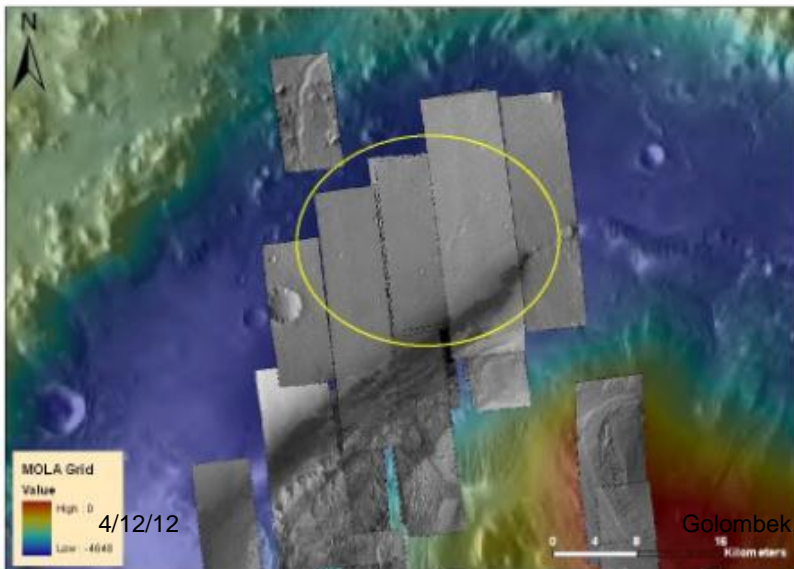
Eberswalde Crater HiRISE



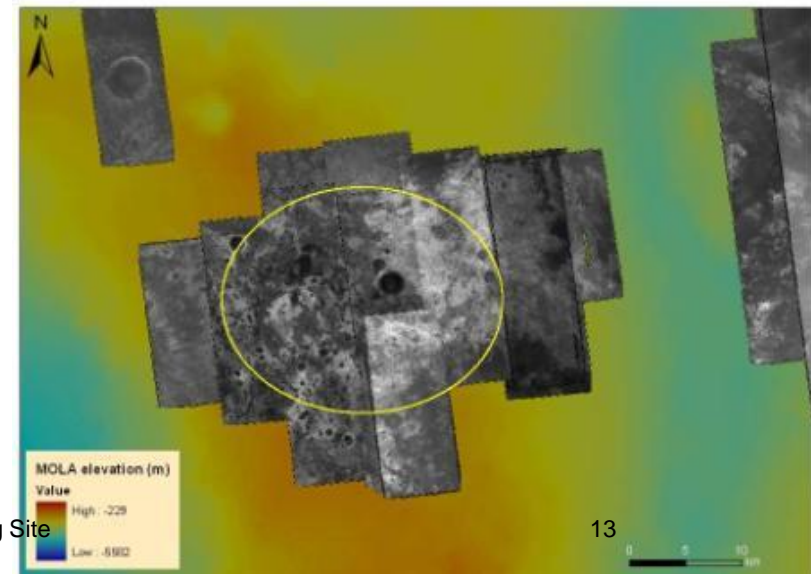
Holden Crater HiRISE



Gale Crater HiRISE

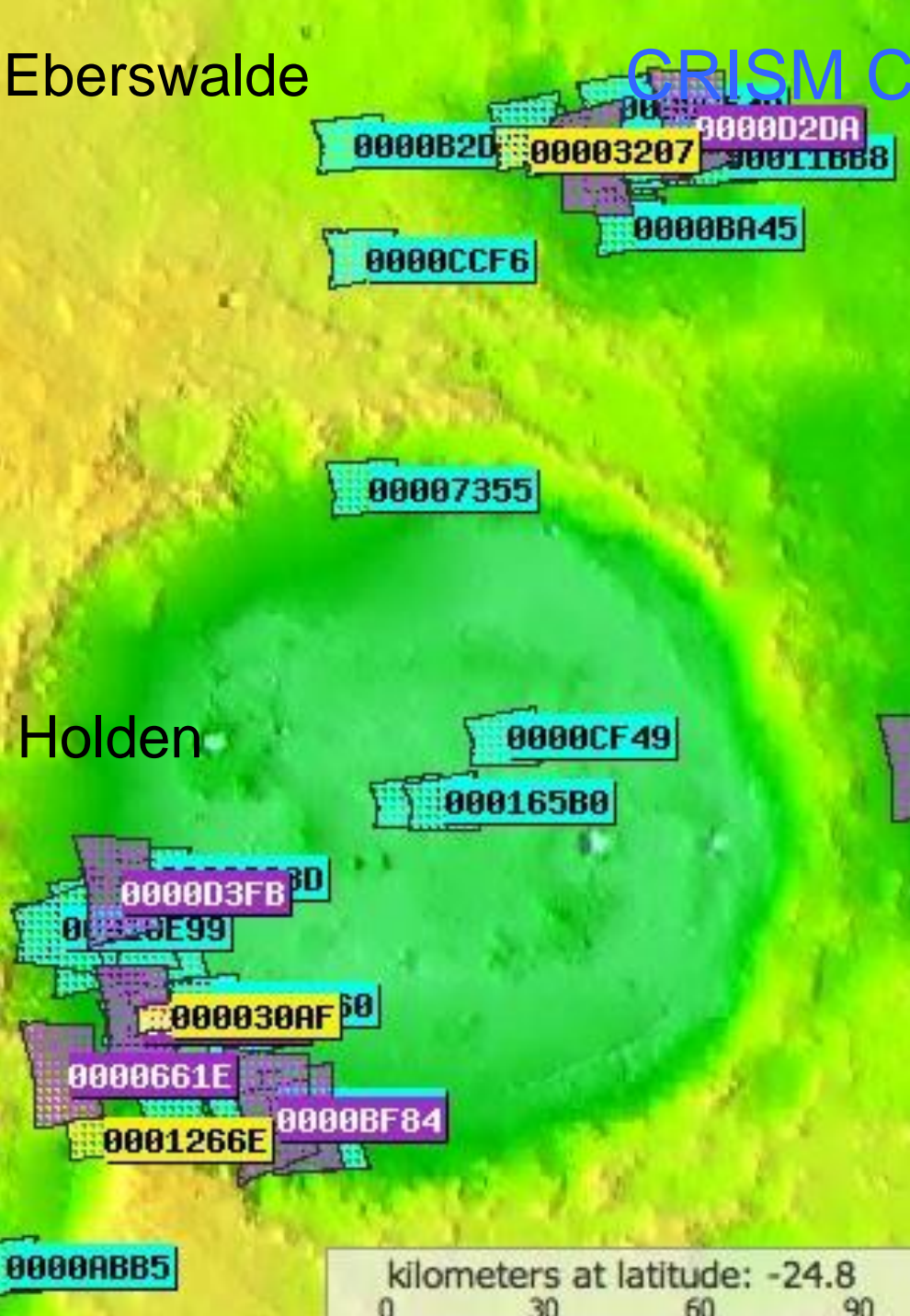


Mawrth Vallis HiRISE





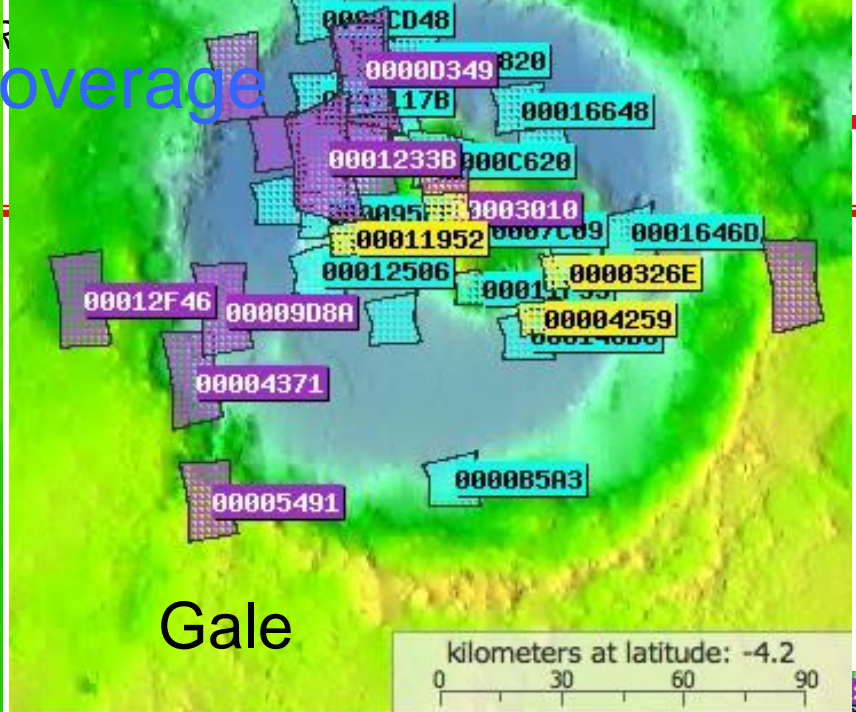
Eberswalde



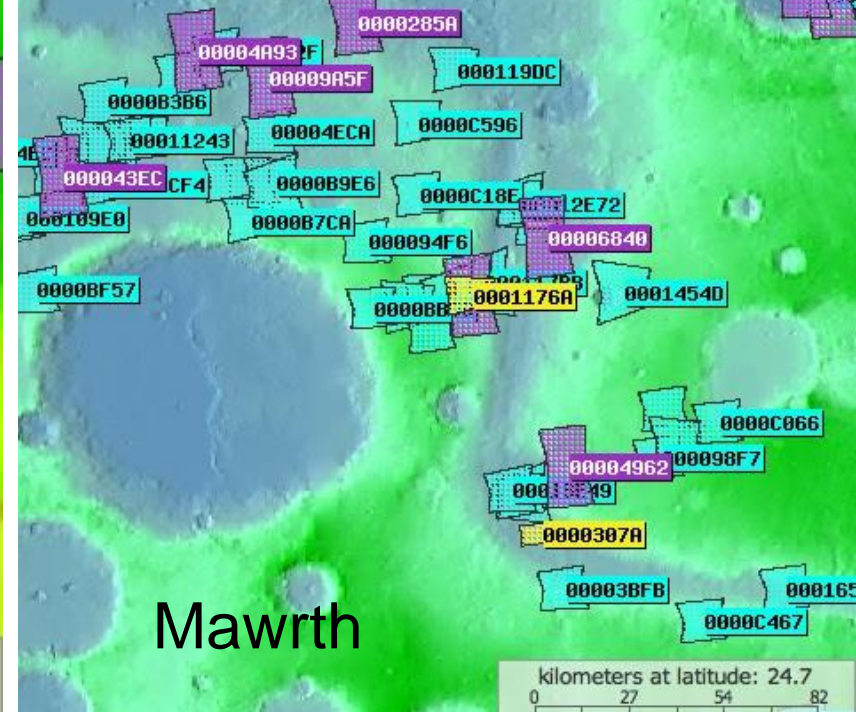
Holden



Gale



Mawrth





# Gale Crater

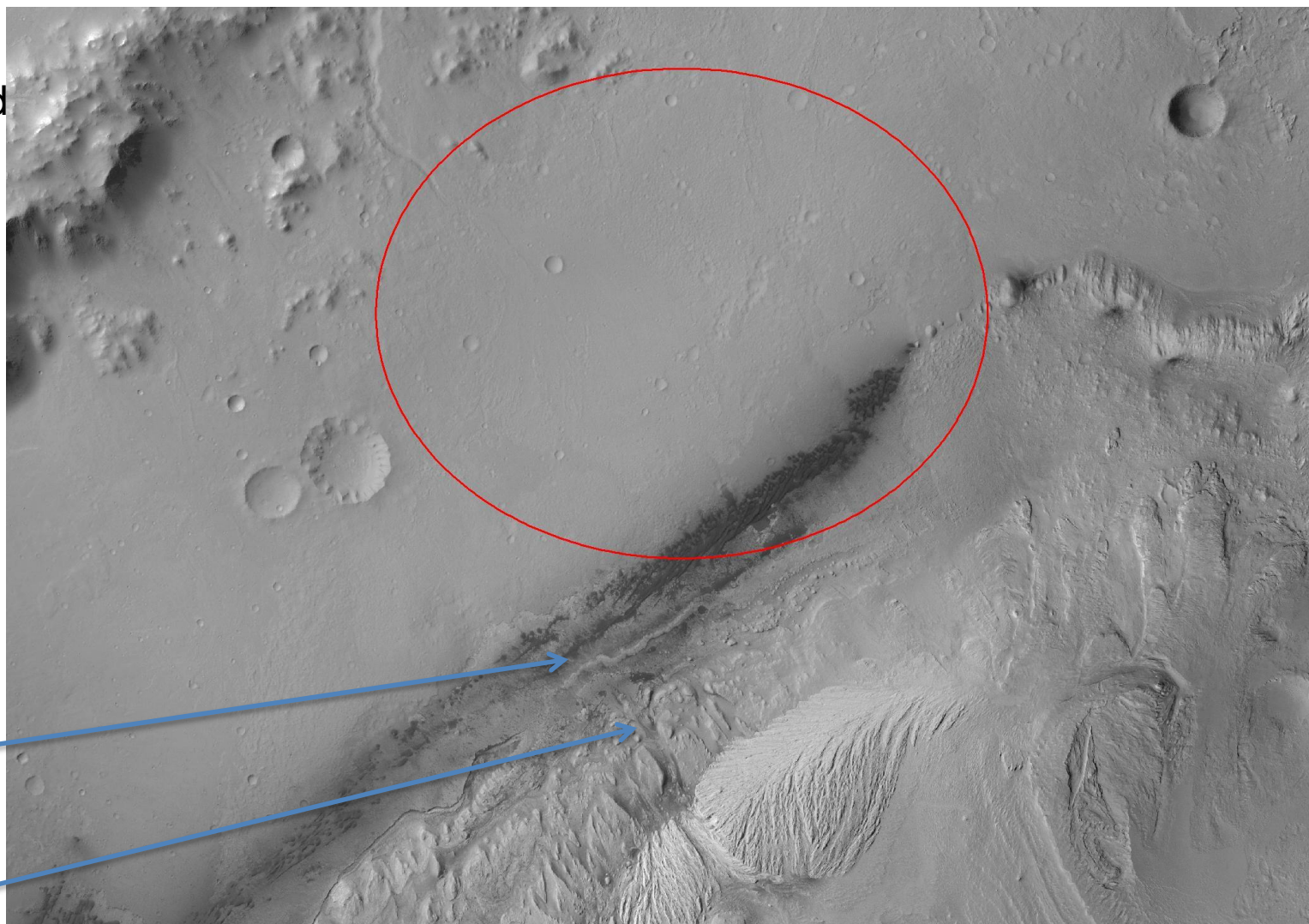
Land on Cratered  
Plains

Smooth, Flat,  
Cratered Plains

“Go To”

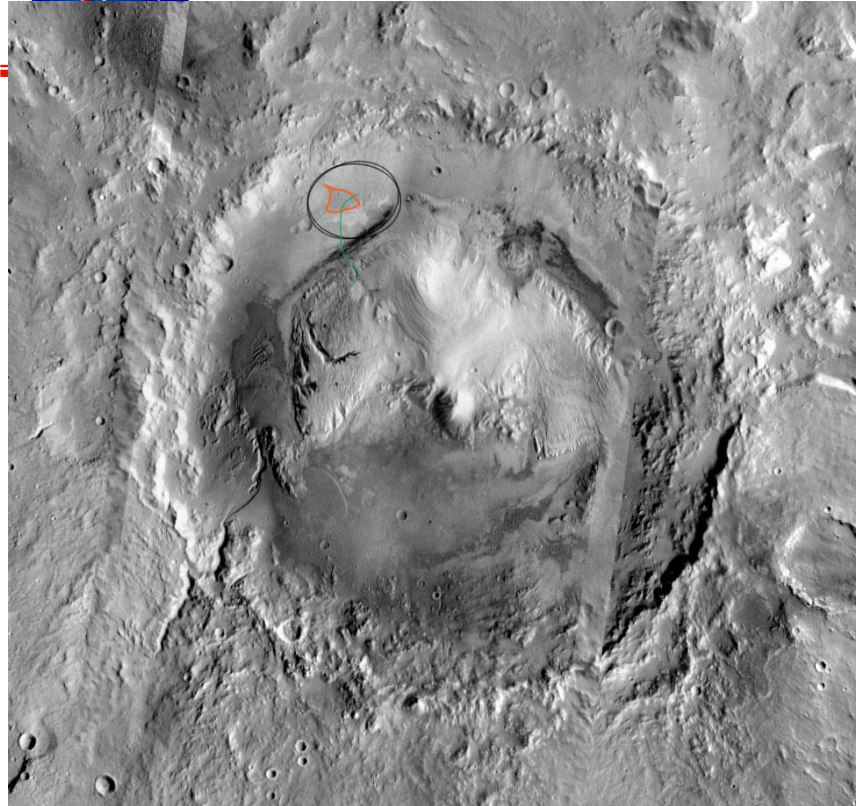
Sample Strata  
Here

Drive up Canyon  
Here

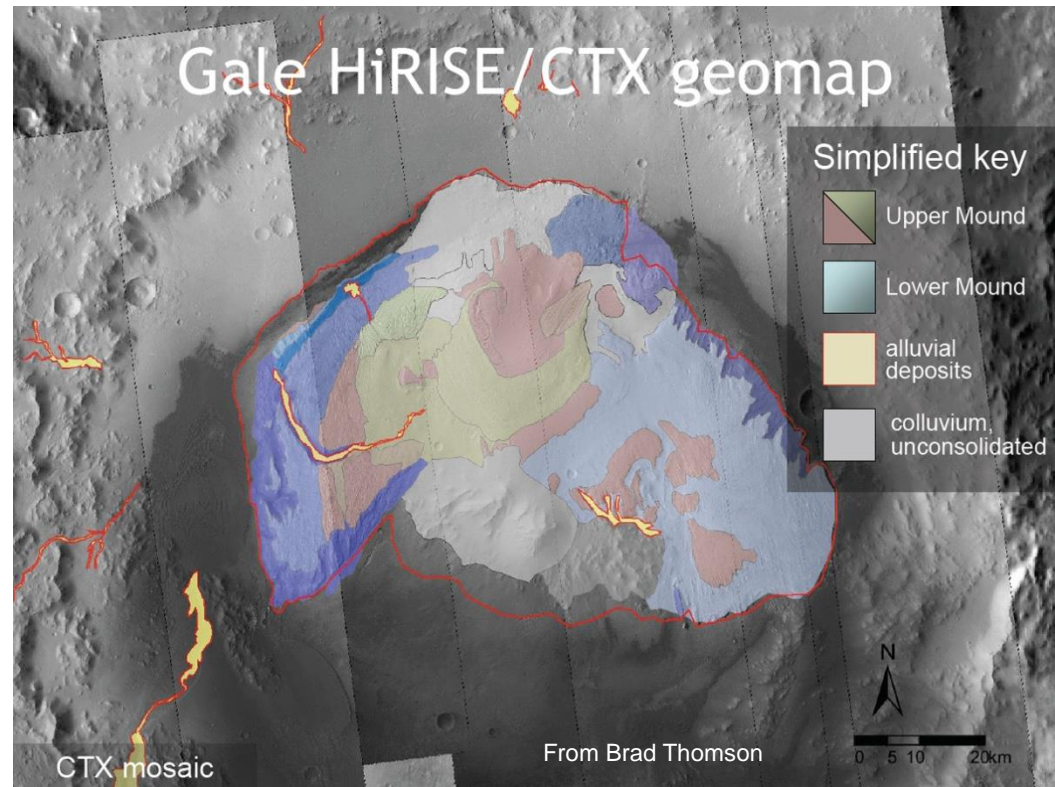
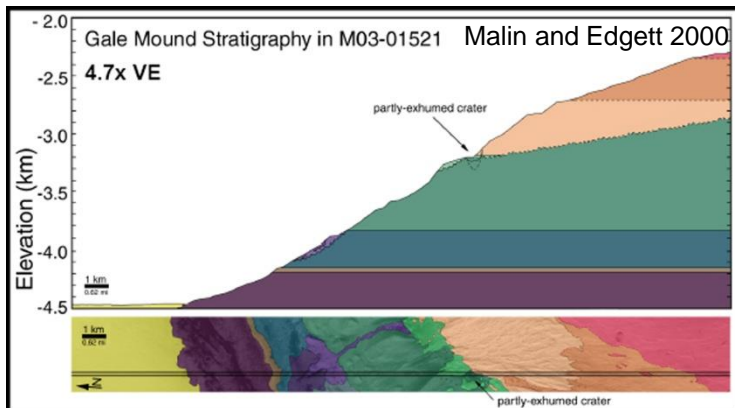




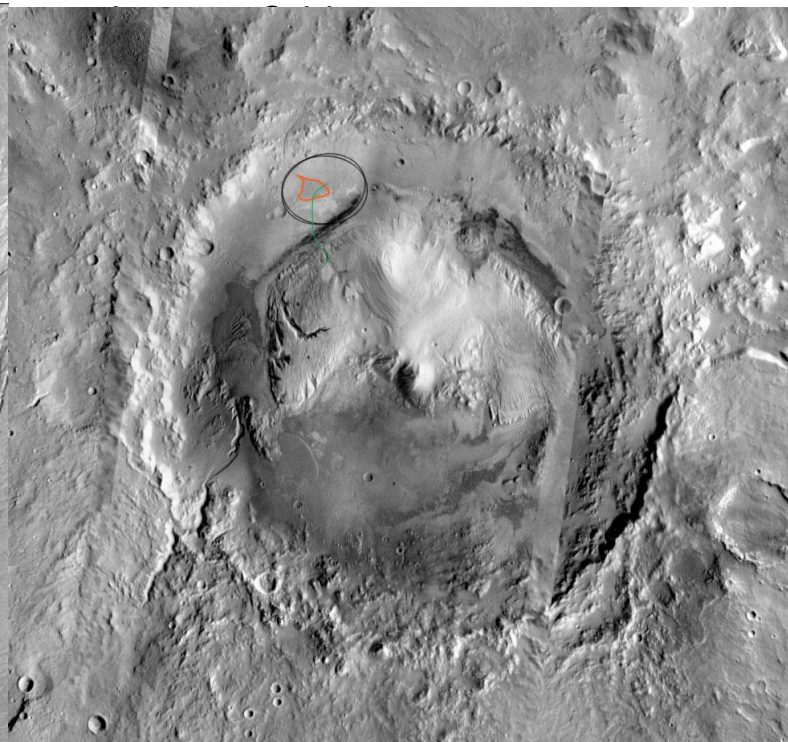
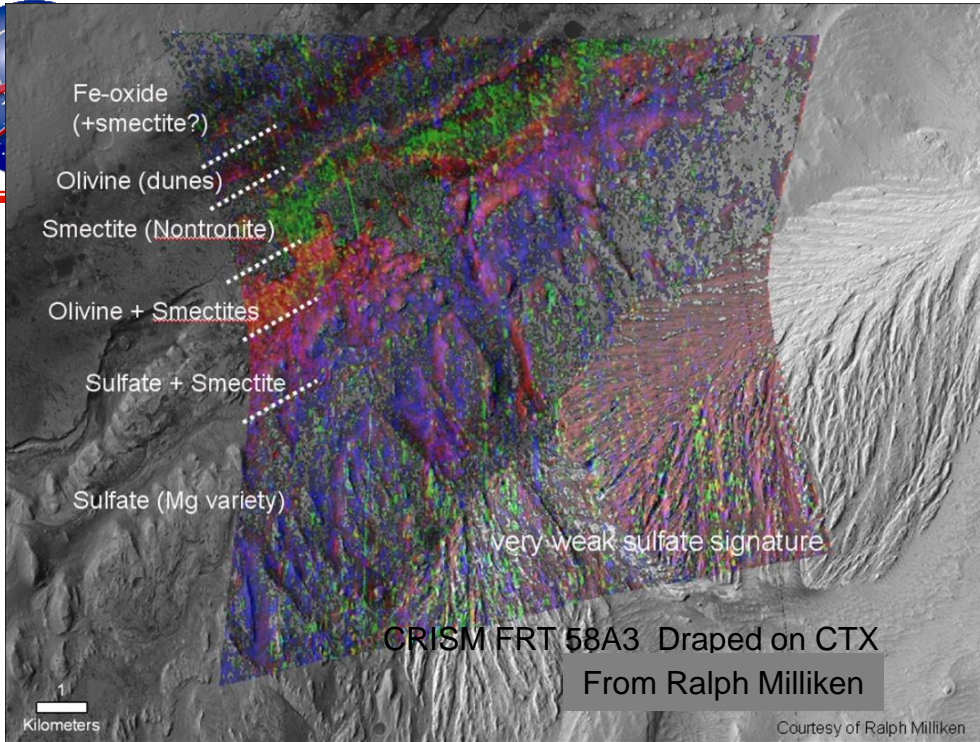
# Mars Reconnaissance Orbiter Gale Crater: K. Edgett, R. Anderson, J. Bell, D. Sumner, R. Milliken



- High diversity of geologic materials with different compositions and depositional conditions
- This diversity is arranged in a stratigraphic context
- Stratigraphy records multiple early Mars environments in sequential order
- Gale is characteristic of a family of craters that were filled, buried, and exhumed, providing insights into an important martian process

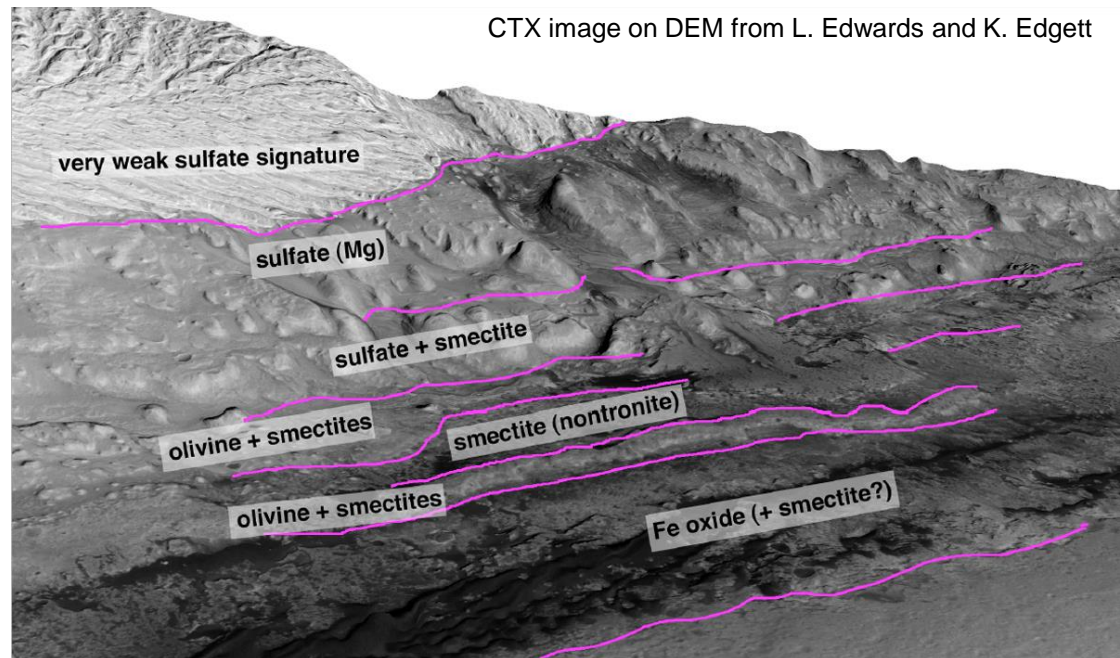




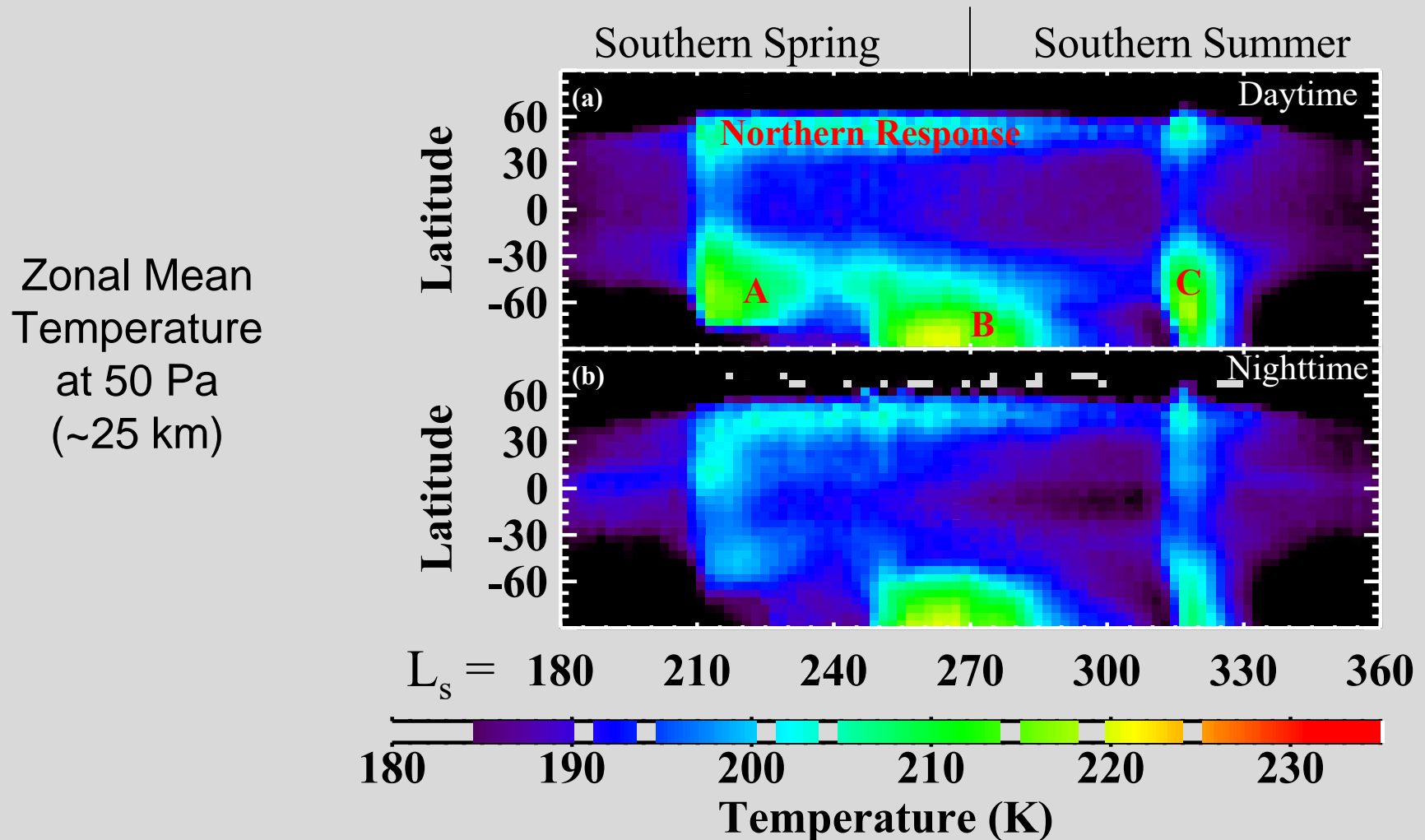


# Gale Crater

CTX image on DEM from L. Edwards and K. Edgett



## Three annually repeating periods of large dust events

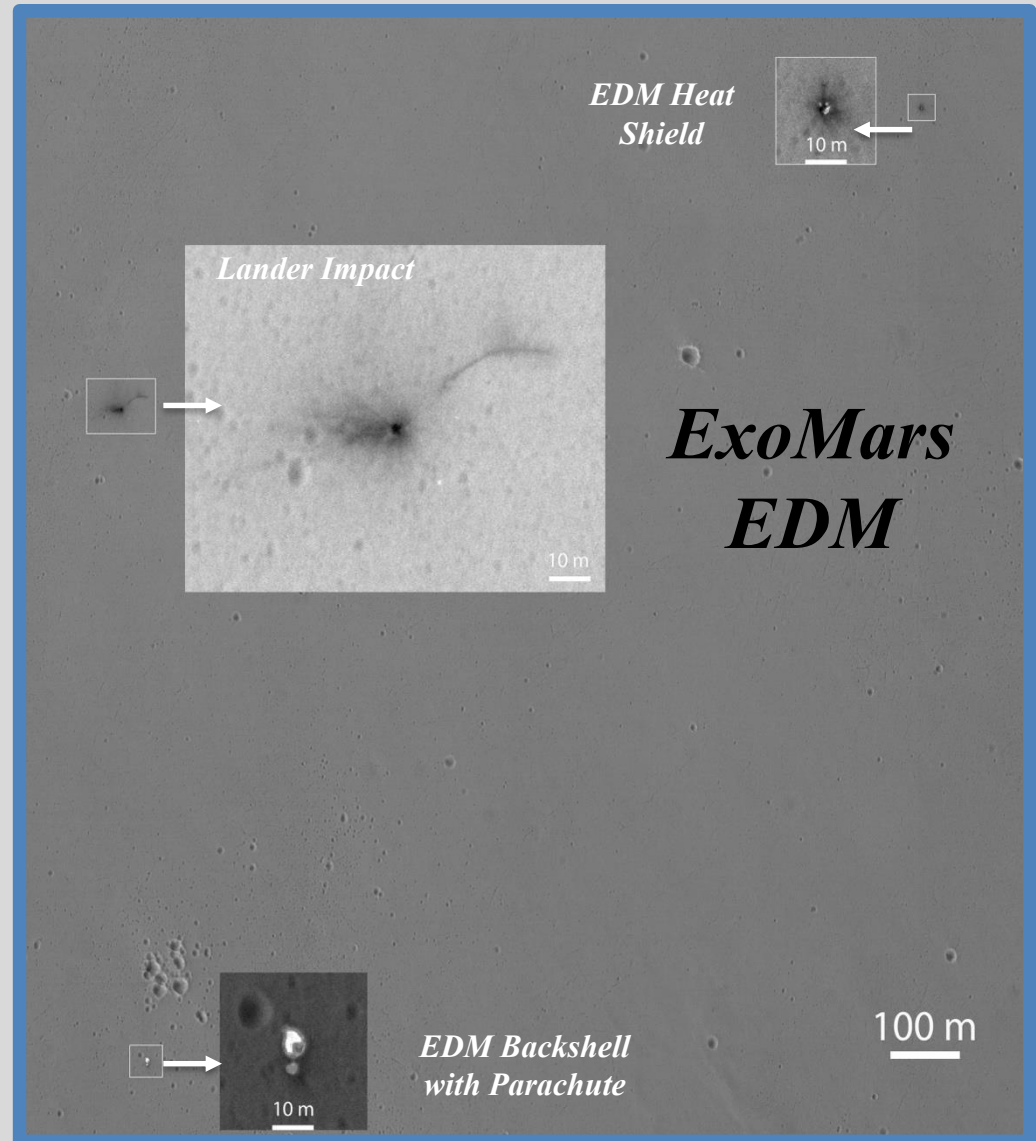
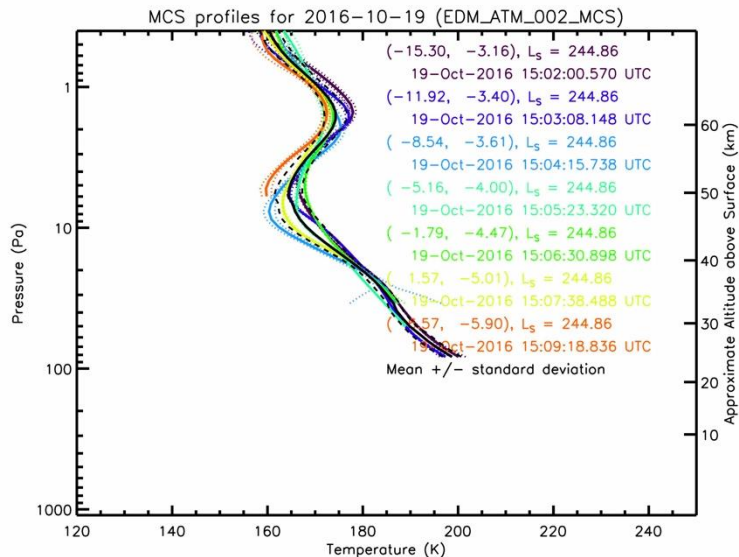


**Reference:** Kass, *et al.* (2016). Interannual similarity in the Martian atmosphere during the dust storm season, *Geophys. Res. Lett.*, 43, doi: 10.1002/2016GRL068978.



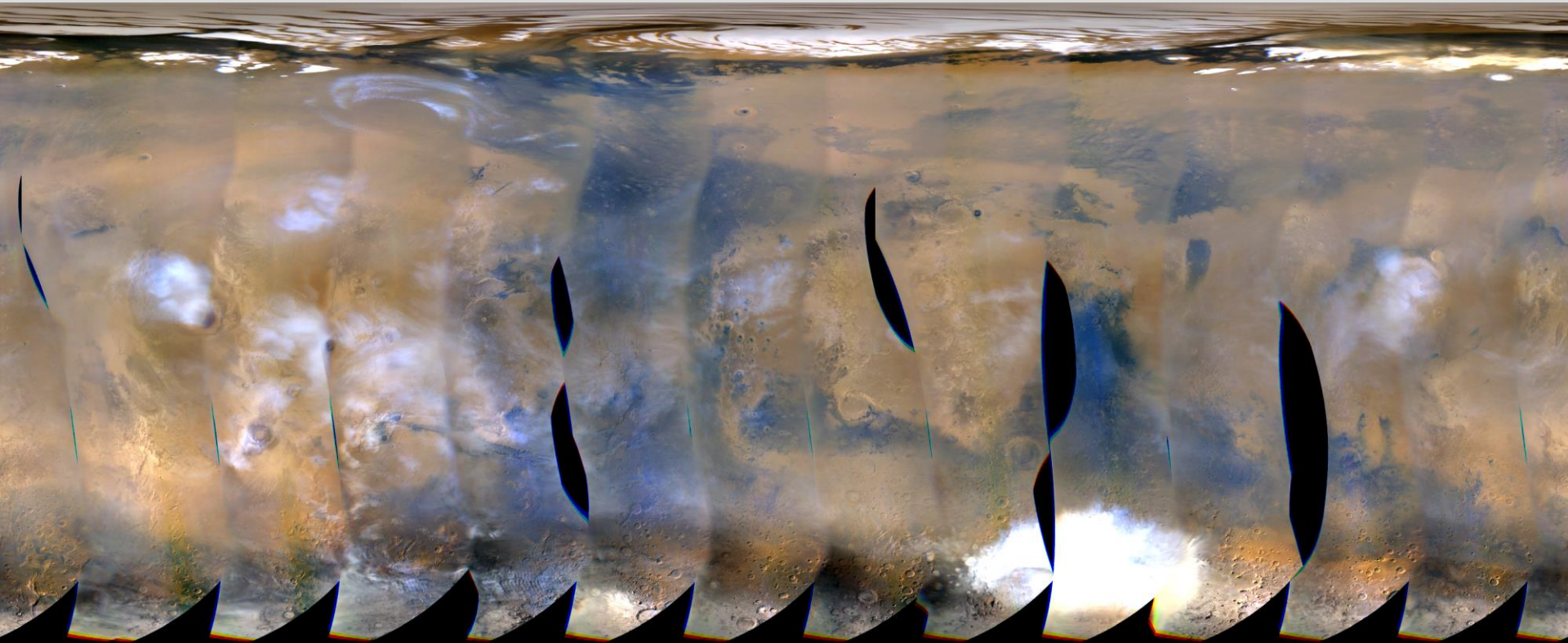
# MRO Supports *ExoMars EDM*

- MCS Profiles the Atmosphere near EDM Entry;
- CTX finds the impact location post-landing;
- HiRISE resolves the EDM flight elements: Lander, Back-shell, Heat Shield.





## Recent MARCI global weather map



March 7, 2016 (  $L_S = 119.1$  )

# Mission Overview



## LAUNCH

- Atlas V 541 vehicle
- Launch Readiness Date: July 2020
- Launch window: July/August 2020

## CRUISE/APPROACH

- ~7 month cruise
- Arrive Feb 2021

## ENTRY, DESCENT & LANDING

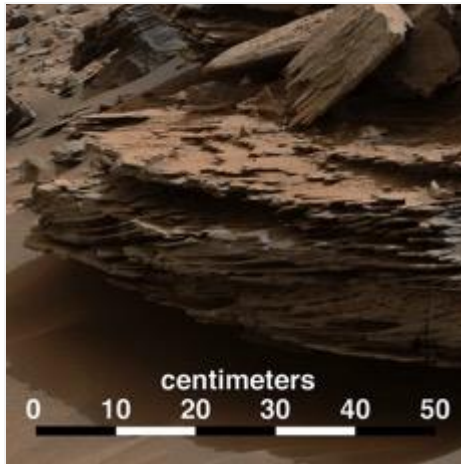
- MSL EDL system (+ [Range Trigger and Terrain Relative Navigation](#)): guided entry and powered descent/Sky Crane
- 16 x 14 km landing ellipse (range trigger baselined)
- Access to landing sites  $\pm 30^\circ$  latitude,  $\leq -0.5$  km elevation
  - Curiosity-class Rover

## SURFACE MISSION

- 20 km traverse distance capability
  - [Enhanced surface productivity](#)
- [Qualified to 1.5 Martian year lifetime](#)
  - Seeking signs of past life
  - Returnable cache of samples
  - Prepare for human exploration of Mars

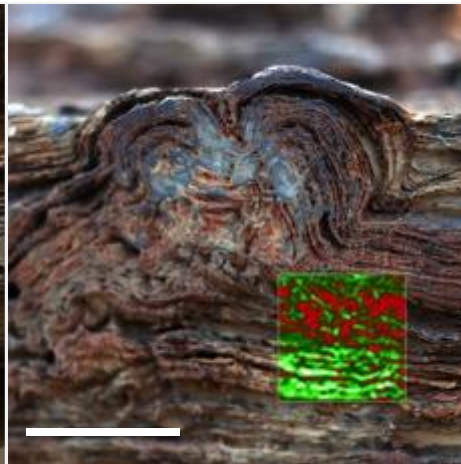


# Mars 2020 Mission Objectives



## GEOLOGIC EXPLORATION

- Explore an ancient environment on Mars
- Understand processes of formation and alteration



## HABITABILITY AND BIOSIGNATURES

- Assess habitability of ancient environment
- Seek evidence of past life
- Select sampling locations with high biosignature preservation potential



## PREPARE A RETURNABLE CACHE

- Capability to collect ~40 samples and blanks, 20 in prime mission
- Include geologic diversity
- Deposit samples on the surface for possible return



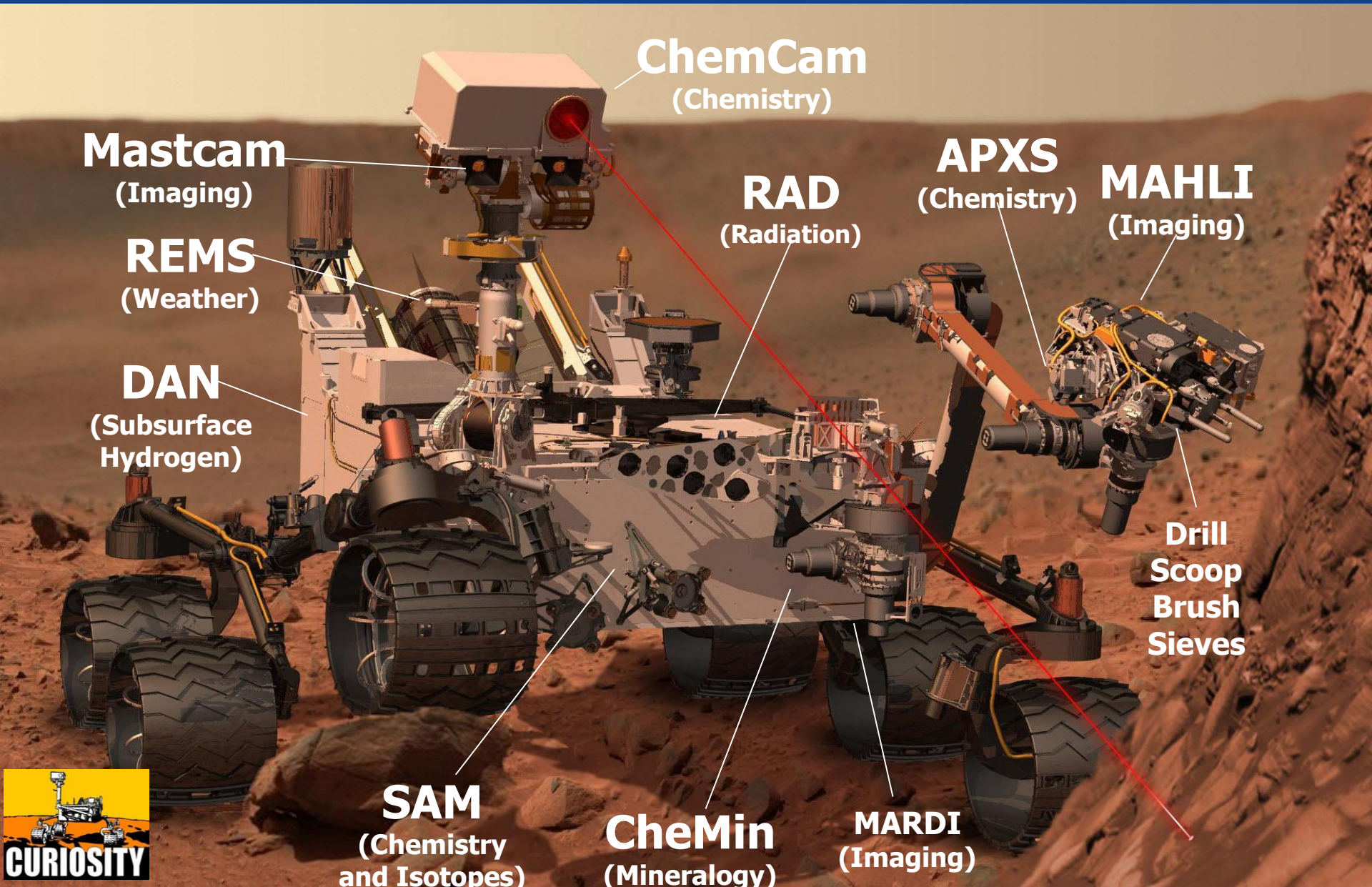
## PREPARE FOR HUMAN EXPLORATION

- Measure temperature, humidity, wind, and dust environment
- Demonstrate In Situ Resource Utilization by converting atmospheric CO<sub>2</sub> to O<sub>2</sub>

# Curiosity science payload



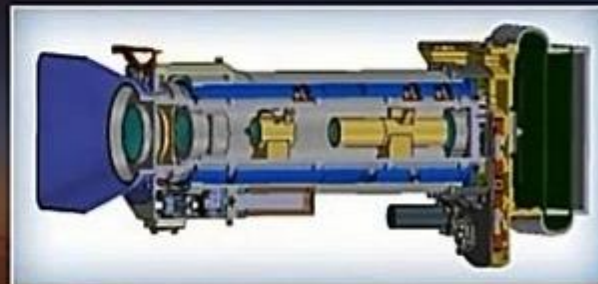
Jet Propulsion Laboratory  
California Institute of Technology





# Mastcam-Z

A Geologic, Stereoscopic, and Multispectral Investigation for  
the NASA Mars-2020 Rover Mission



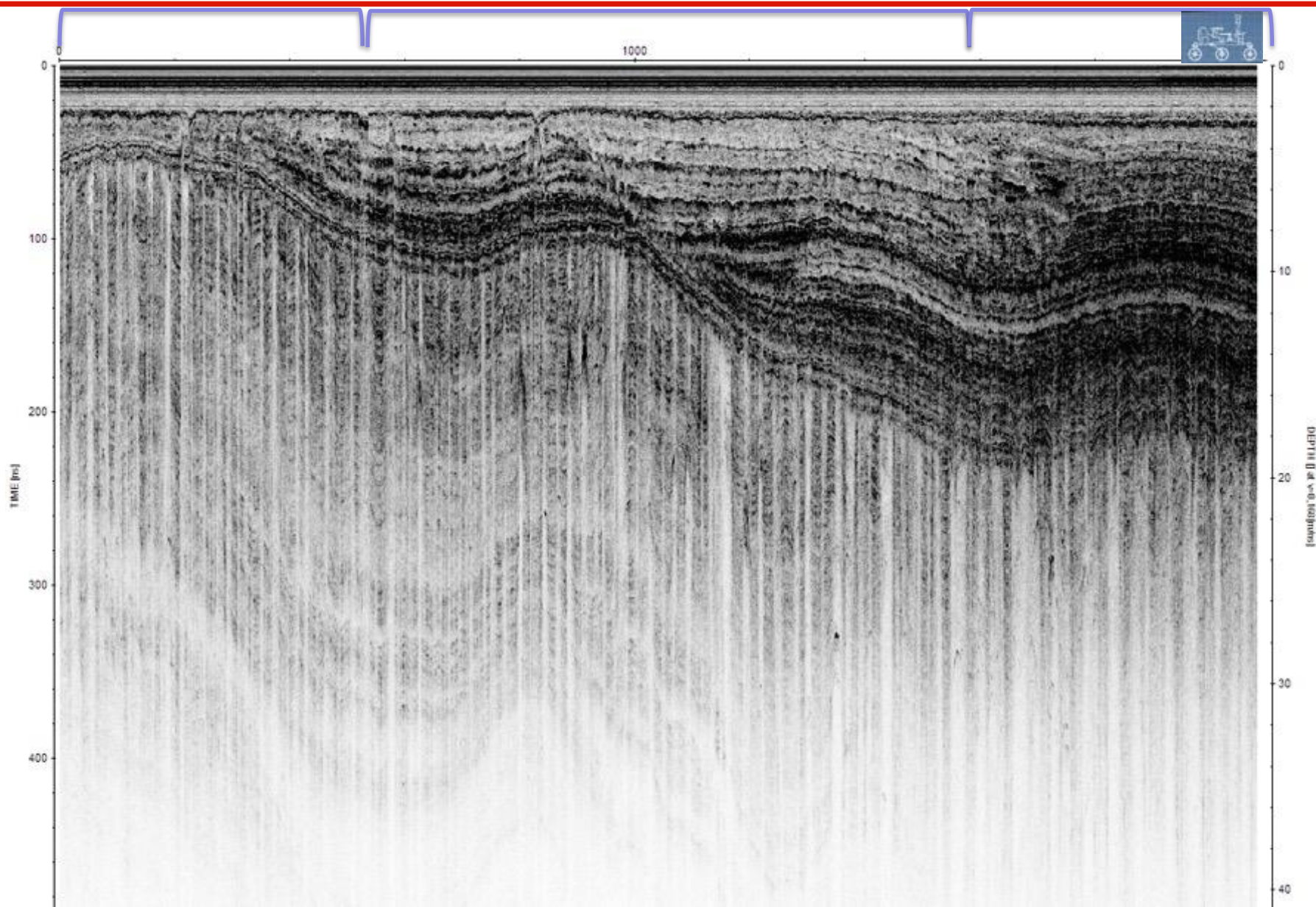
improved stereo zoom camera

# RIMFAX: A View Beneath the Surface

Drive 1

Drive 2

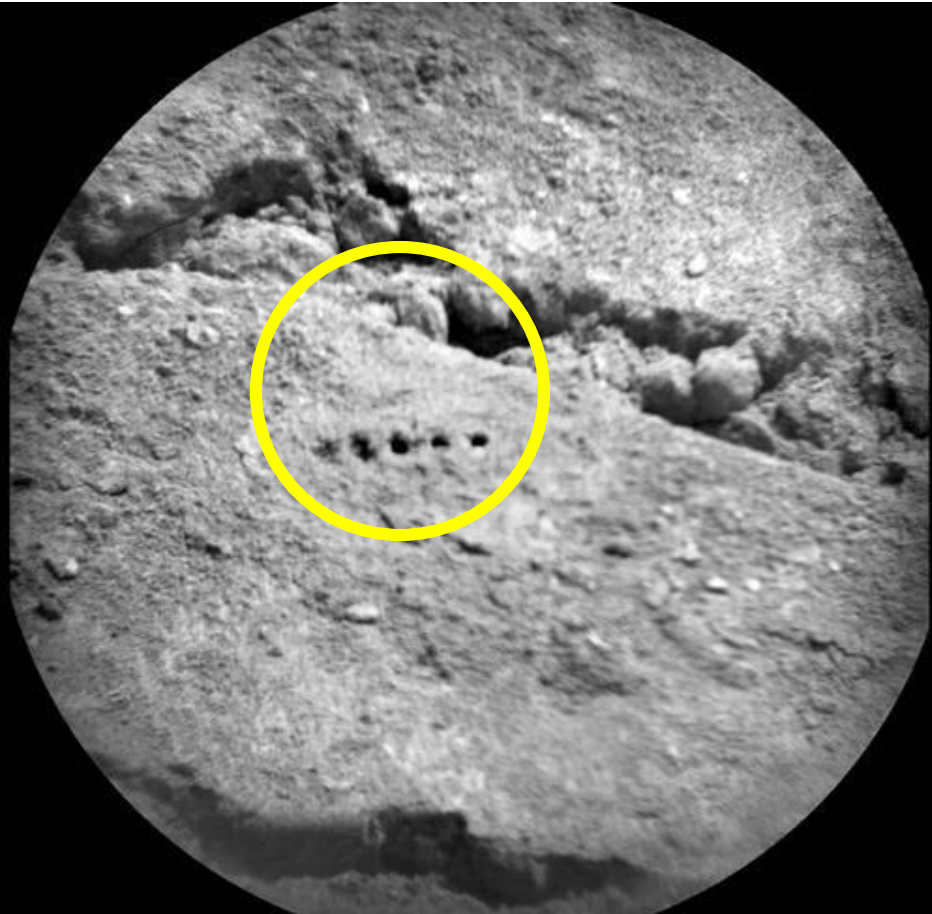
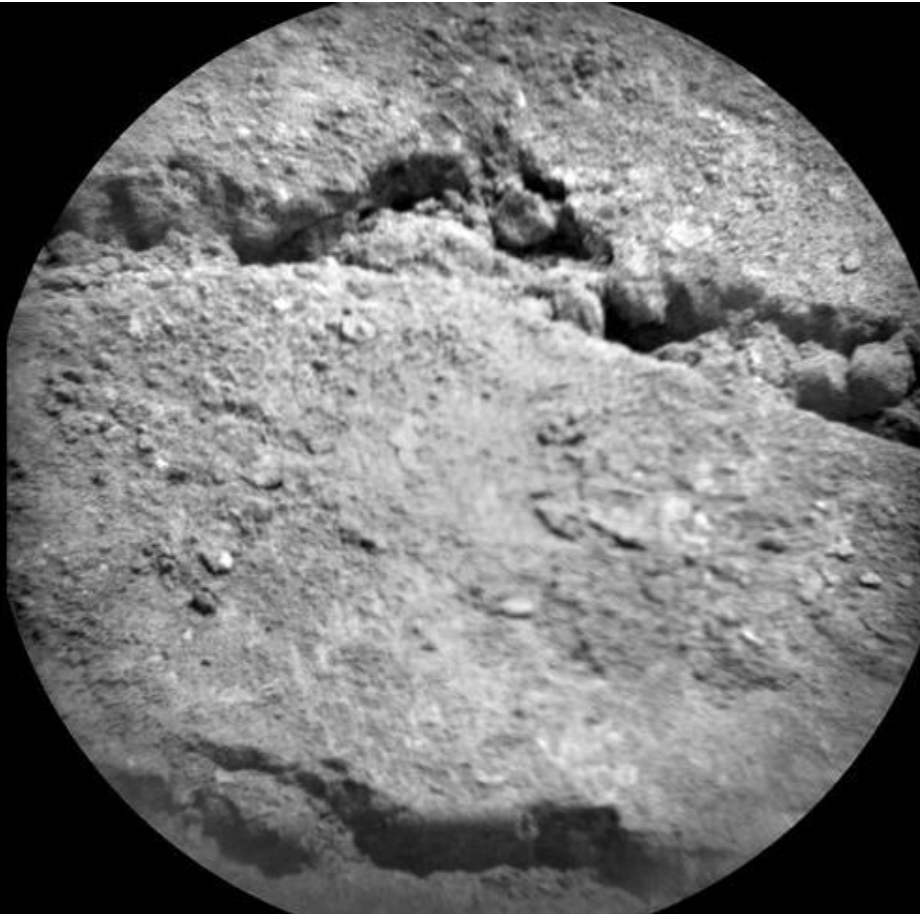
Drive 3



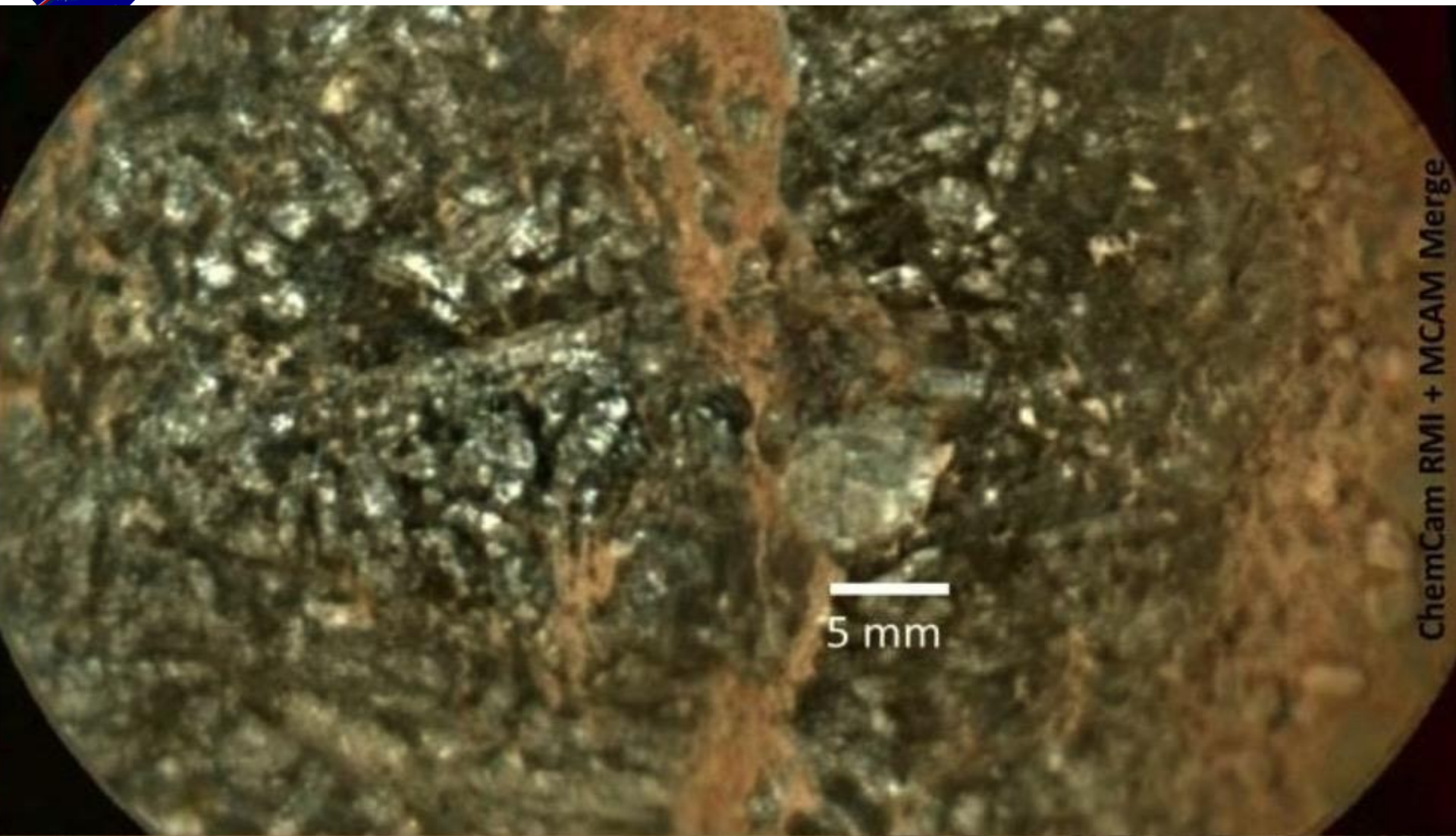


Before

After



ChemCam Target: Beechey (Sol 19)  
Power: 1 Gigawatt  
5-spot raster, shots per spot: 50







2 mm

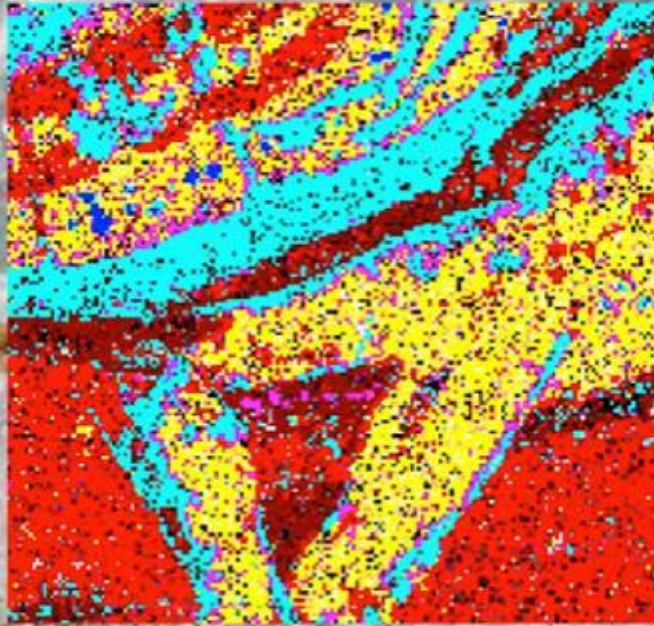


WATSON

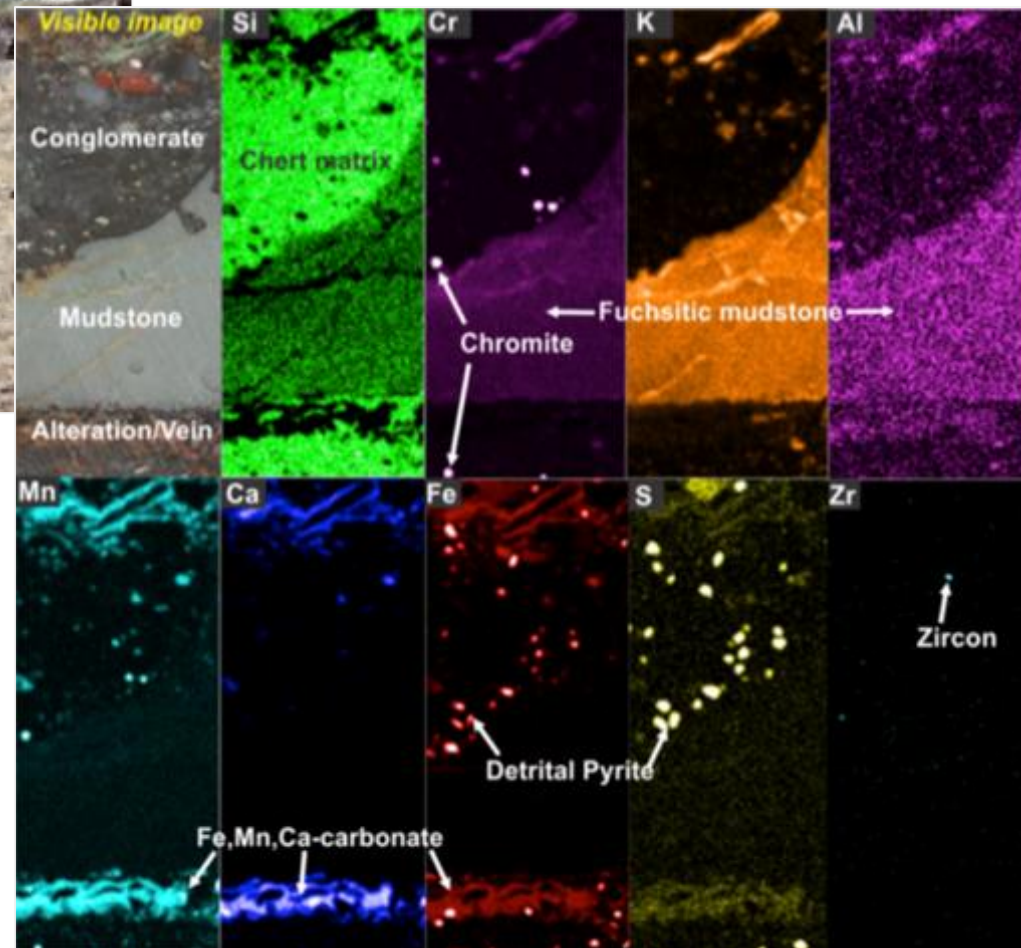




# PIXL: Planetary Instrument For X-ray Lithochemistry



micro x-ray fluorescence  
for elemental mapping  
with sub-mm resolution



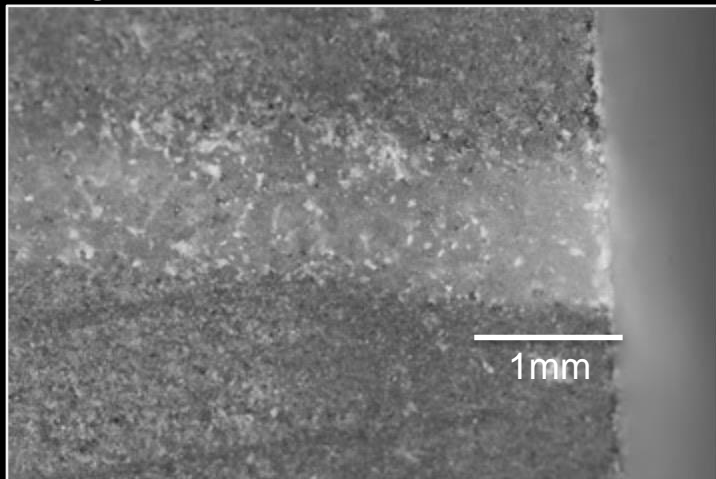
# SHERLOC: Scanning Habitable Environments with Raman & Luminescence for Organics & Chemicals

SHERLOC's  
view through  
WATSON

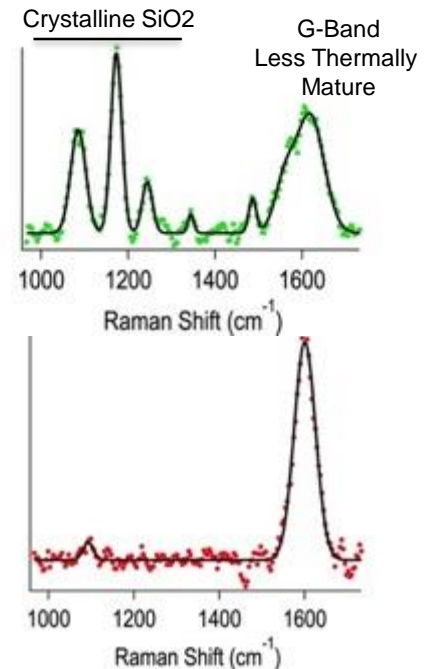
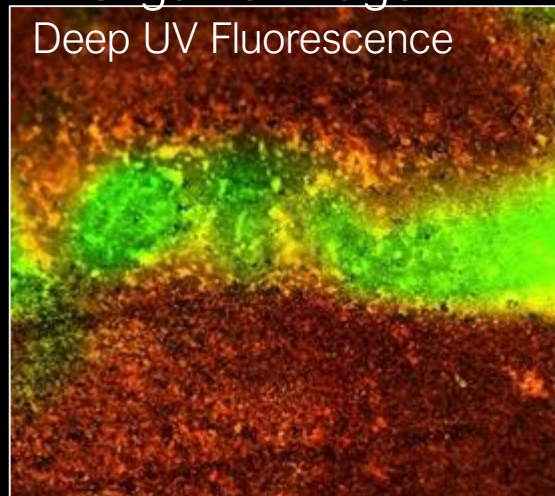


Organic & Mineral  
Analyzer  
Deep UV Raman

High Resolution Camera



Organic Image  
Deep UV Fluorescence



More Mature

Less Mature

Organic Maturity

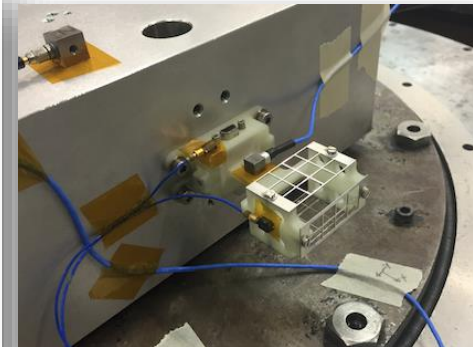
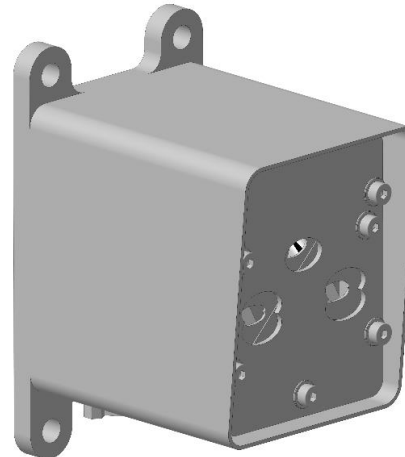
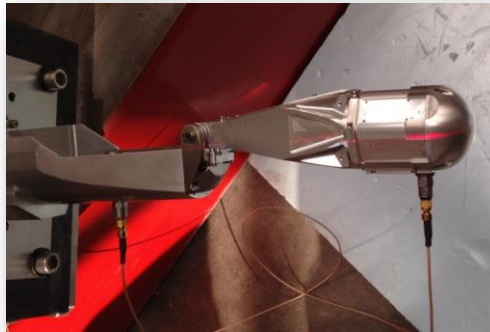
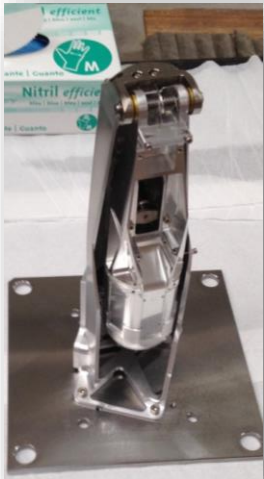
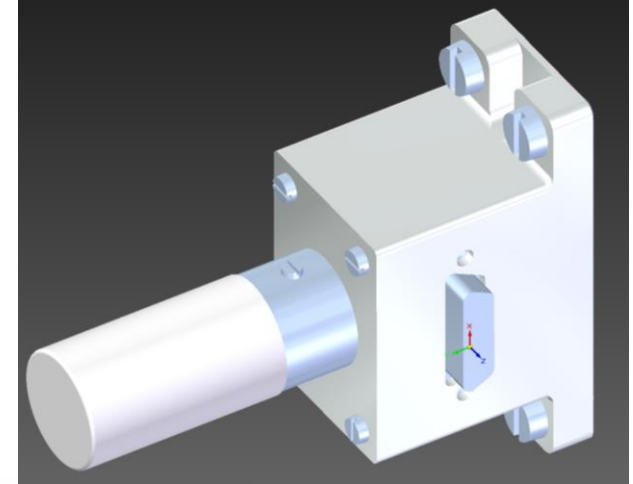
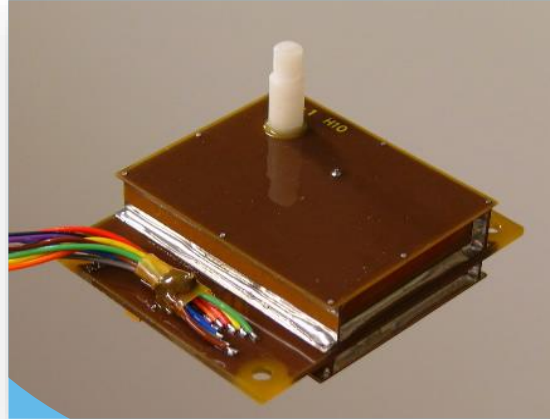


## Mars Environmental Dynamics Analyzer

**A. Folded configuration**

Labels in the diagram include: RSM, RS+Dust Sensor, Wind Sensor, Deployment System, ICU+PS (inside), Air Temp Sensor, Humidity Sensor, Thermal Infrared Sensor, and E.

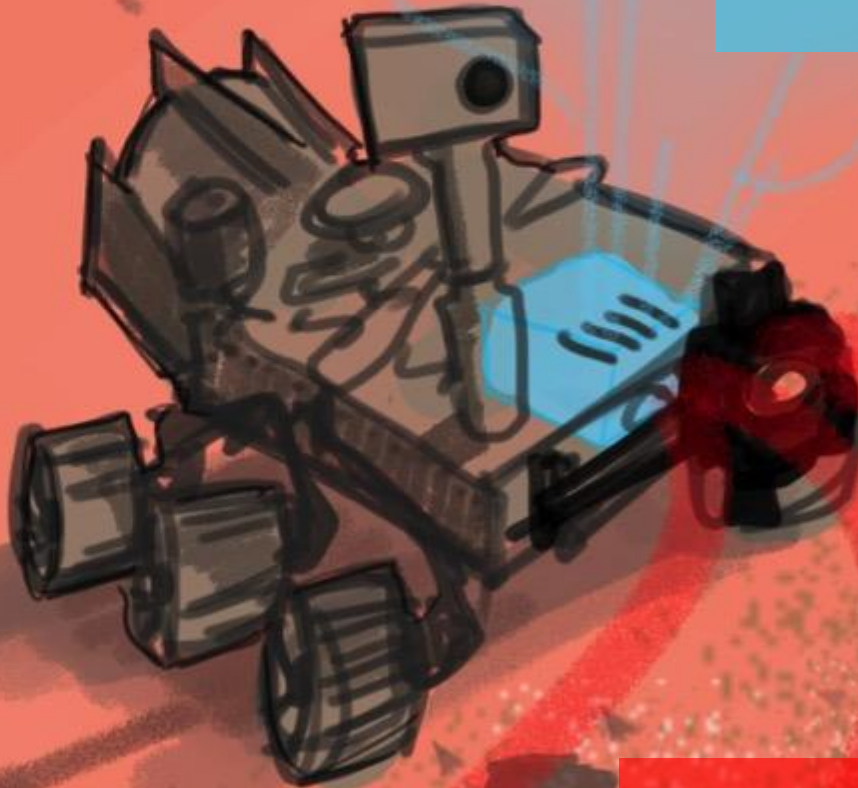
# MEDA subsystems





MOXIE

STEP 02:  
OXYGEN



STEP 01:  
MARS ATMOSPHERE

# Spacecraft Build Approach



## Launch Vehicle

- KSC/Launch Services Program procurement

## MMRTG

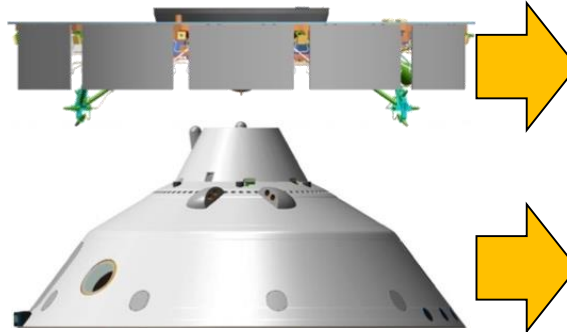
- DoE procurement to industry

## Science & Exploration Technology Investigations

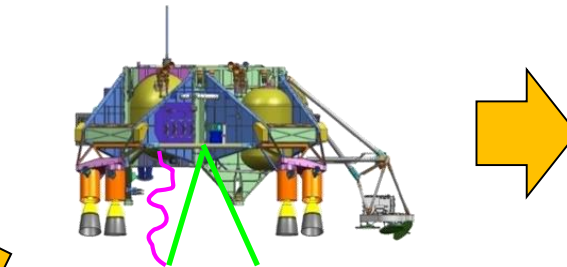
- Source per proposals via AO selection

## MEDLI2

- NASA Centers (LaRC, ARC, and JPL)



- Built in-house at JPL
- Lowest cost and risk per make-buy study and industry RFIs



- Built by Lockheed-Martin/Denver
- Procure as sole source—most cost effective

- Built in-house at JPL
- Major industry subcontracts/components
- Rebuild in-house due to criticality of EDL and rover interface



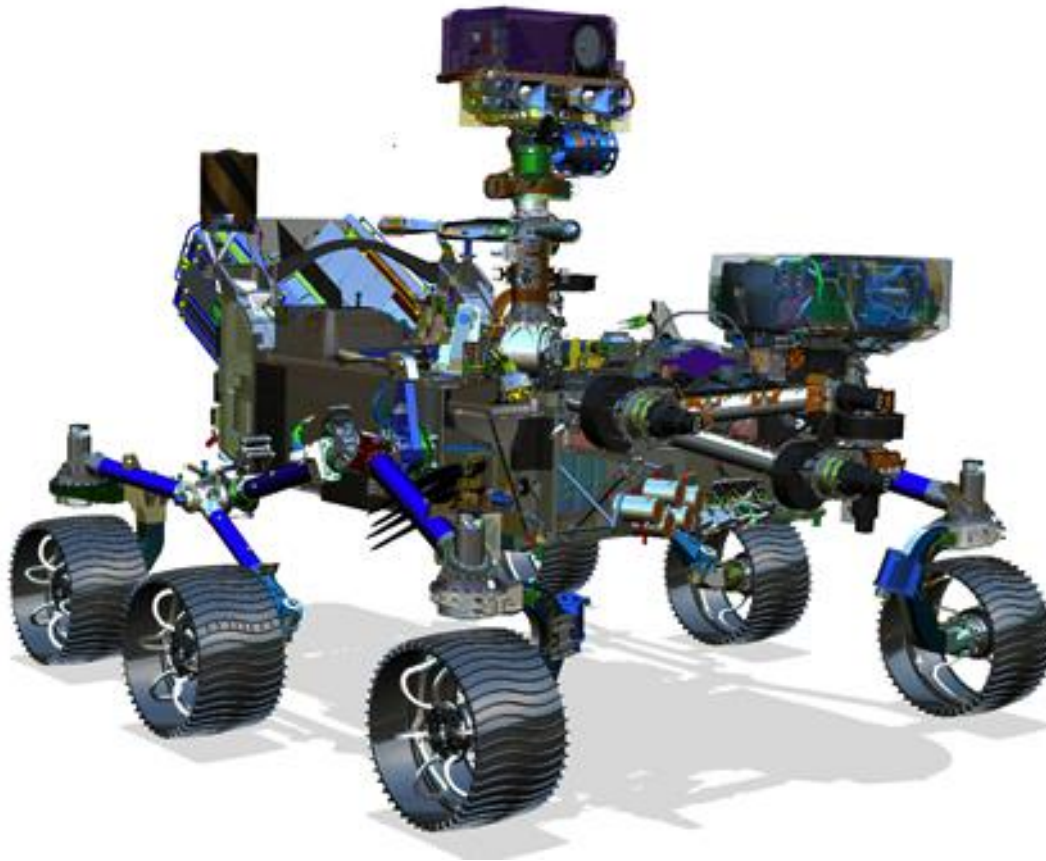
- Built in-house at JPL
- Major industry subcontracts/components
- Spanish contributed High Gain Antenna
- Rebuild in-house due to complexity of vehicle, residual hardware, criticality of EDL and rover interface, operations experience



- Built by Lockheed-Martin/Denver
- Procure as sole source—most cost effective



# Mars 2020 Rover Concept



## High Heritage from MSL

- Avionics
- Power
- GN&C
- Telecom
- Thermal
- Mobility

## Changed

- New Science Instrument Suite
- New Sampling Caching System
  - Modified Chassis
  - Modified Rover Harness
  - Modified Surface FSW
- Modified Rover Motor Controller
  - Modified Wheels

# Mars 2020 Mission Objectives

## ■ Conduct Rigorous *In Situ* Science

- A. **Geologic Context and History** Carry out an integrated set of context, contact, and spatially-coordinated measurements to characterize the geology of the landing site
- B. **In Situ Astrobiology** Using the geologic context as a foundation, find and characterize **ancient** habitable environments, identify rocks with the highest chance of preserving **signs of ancient** Martian life if it were present, and within those environments, seek the signs of life

## ■ Enable the Future

- C. **Sample Return** Assemble rigorously documented and returnable cached samples for possible return to Earth
- D. **Human Exploration** Facilitate future human exploration by making significant progress towards filling major strategic knowledge gaps and...

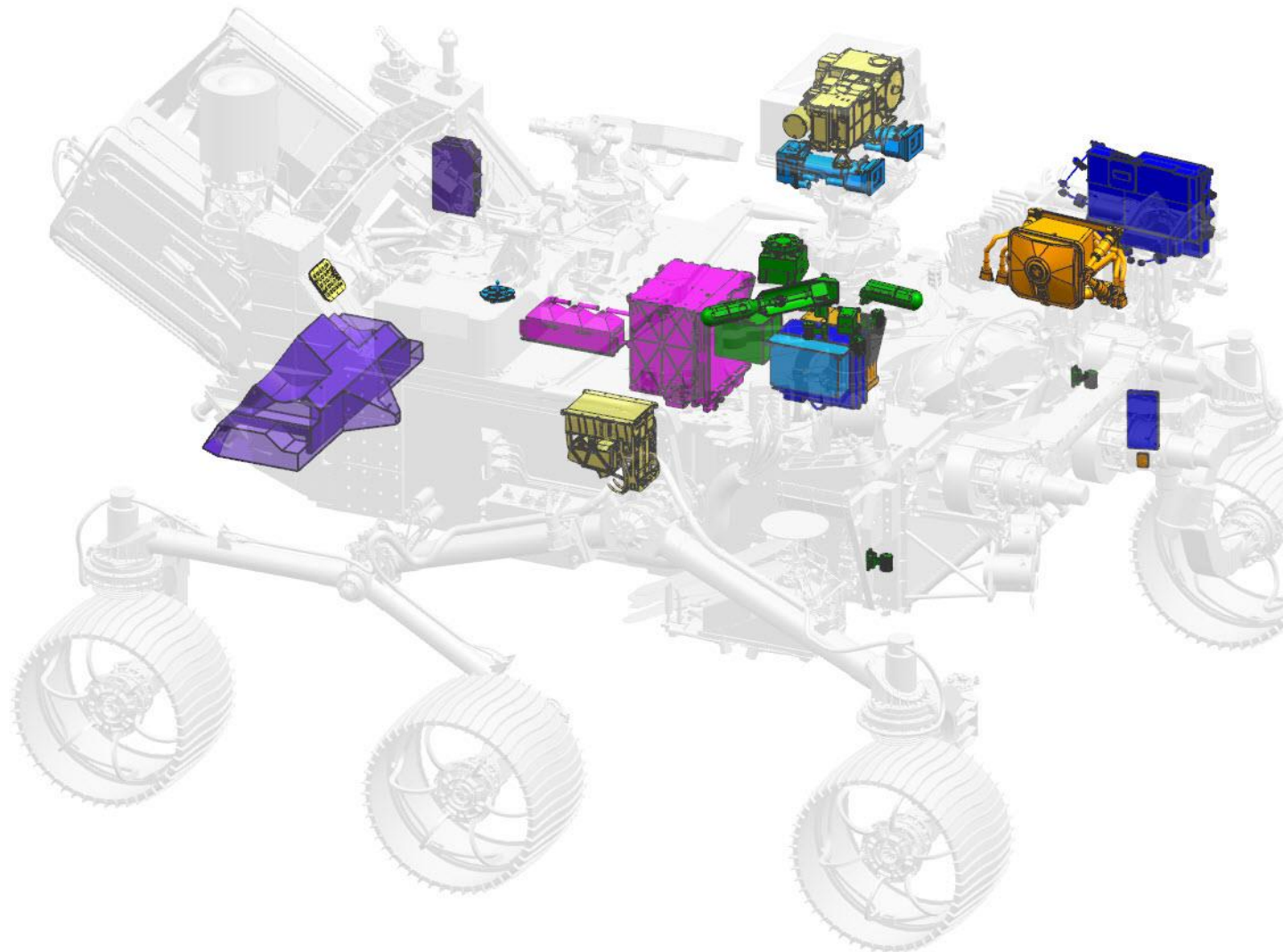
**Technology** ...demonstrate technology required for future Mars exploration

## ■ Execute Within Current Financial Realities

- Utilize MSL-heritage design and a moderate instrument suite to stay within the resource constraints specified by NASA

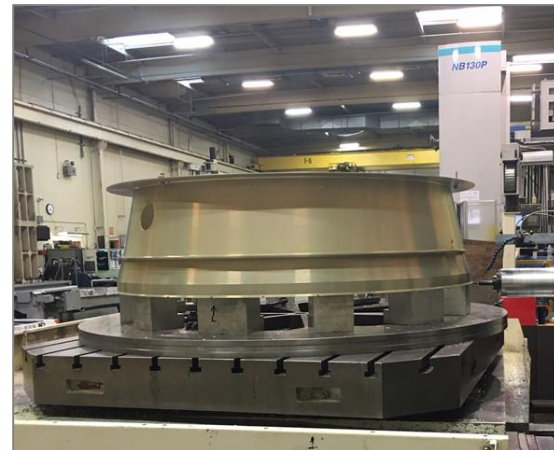
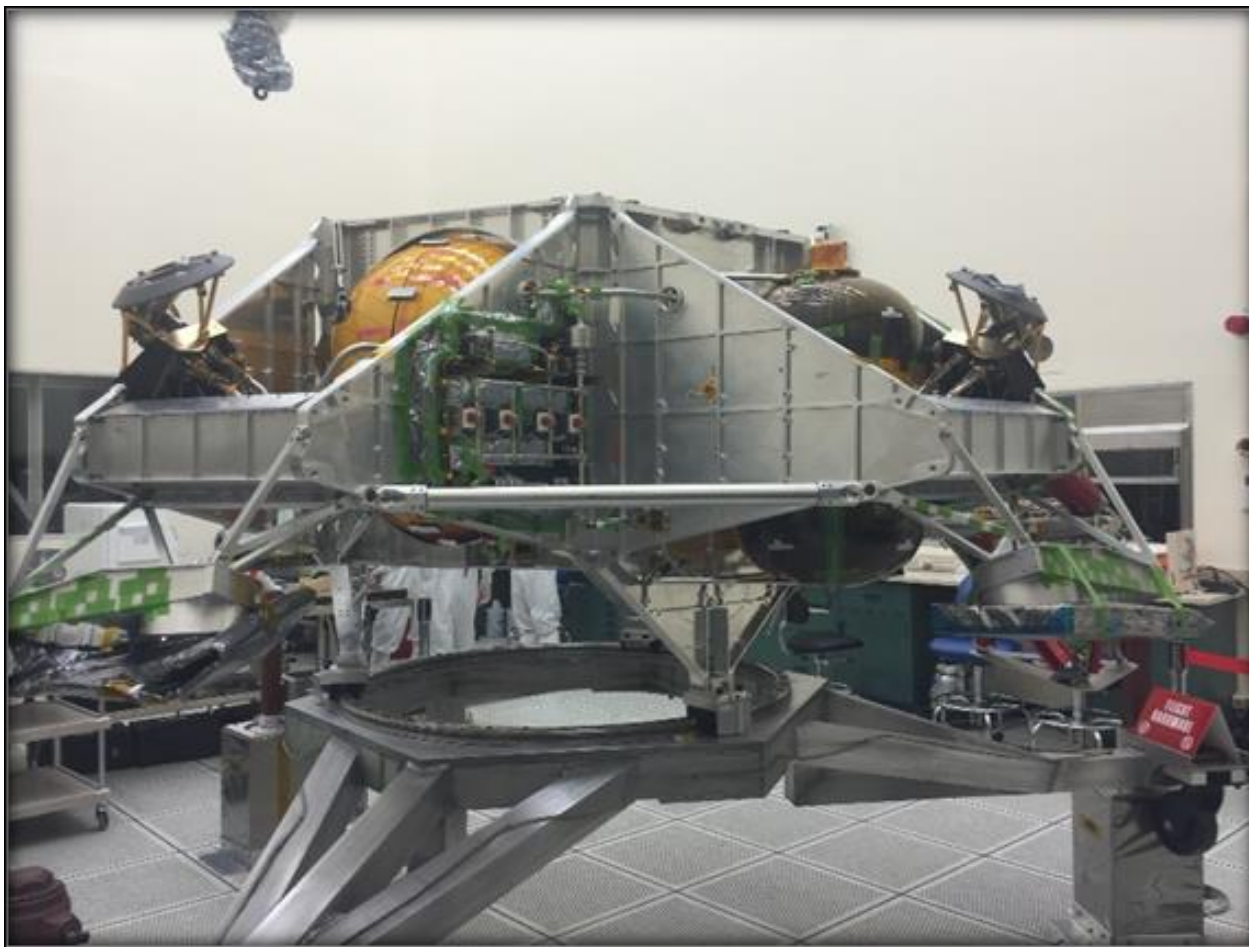


# Mars 2020 Payload Family Picture



Instrument Key
<b>Mastcam-Z</b> Stereo Imager
<b>MEDA</b> Mars Environmental Measurement
<b>MOXIE</b> In-Situ Oxygen Production
<b>PIXL</b> Microfocus X-ray fluorescence spectrometer
<b>RIMFAX</b> Ground Penetrating Radar
<b>SHERLOC</b> Fluorescence and Raman spectrometer and Visible context imaging
<b>SuperCam</b> LIBS and Raman

# Cruise / EDL Systems – In Assembly

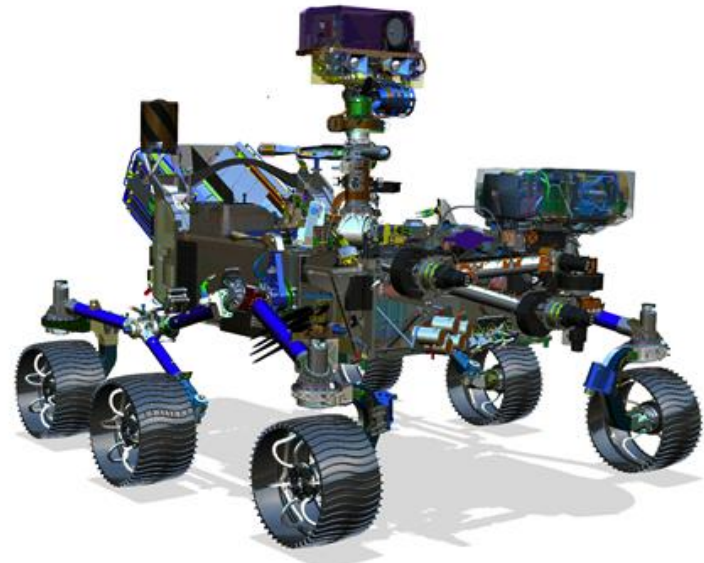




# Project Overview

## *Salient Features*

- *Category: 1*
  - *Risk Class: A-tailored*
- *Directed, JPL in-house implementation*
  - *High heritage MSL design*
- *Modifications only as necessary to accommodate new payload and Sampling / Caching System (SCS)*
- *Planetary Protection Category V Restricted Earth Return per Level 1 Requirements*



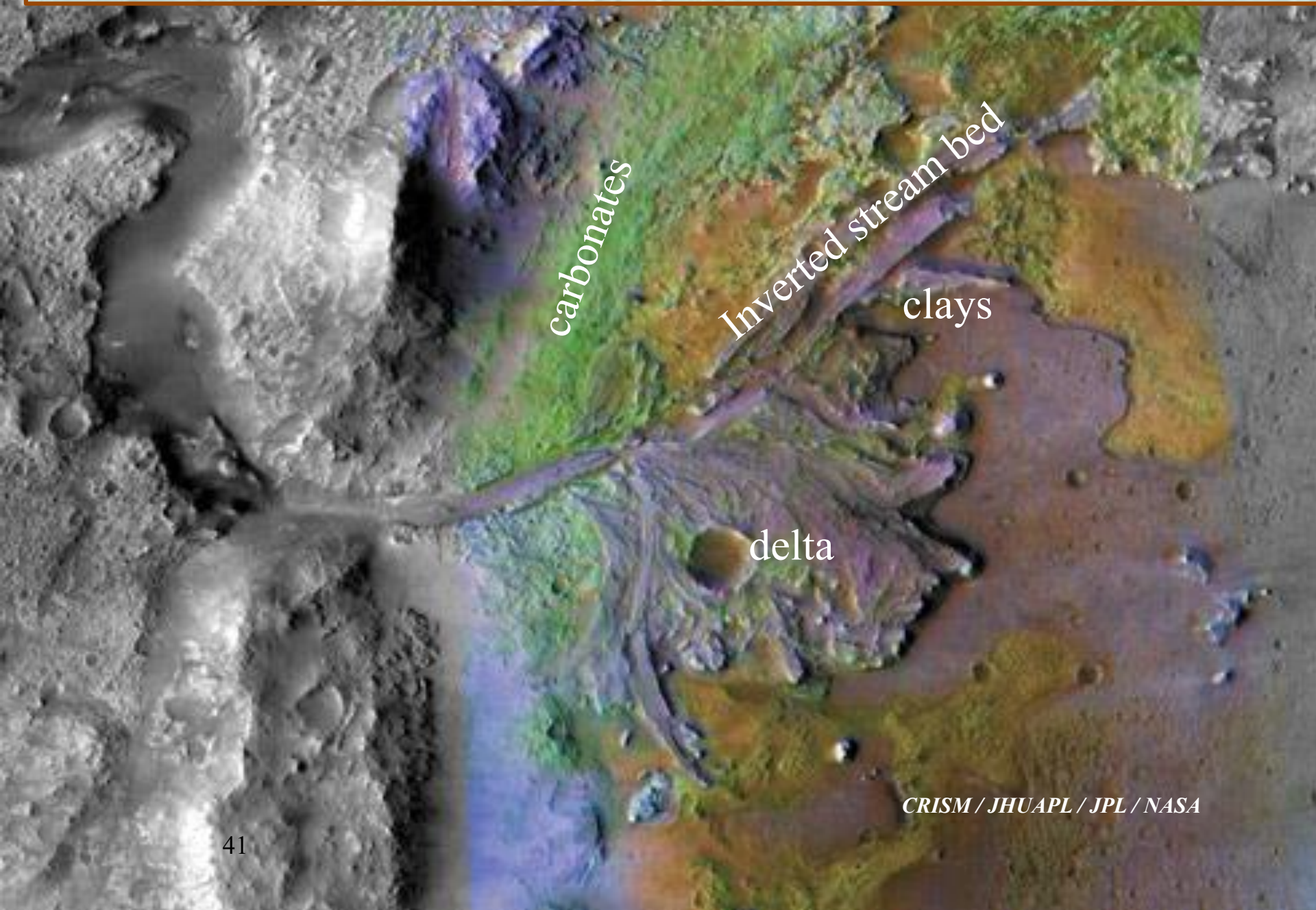
## *Science*

- *Assess past habitability of an astrobiologically relevant ancient environment on Mars*
- *Assess biosignature preservation potential with the environment and search for biosignatures*
  - *Assemble cached samples for possible future return to Earth*

## *Technology*

- *Advance technologies with applications to future human and robotic explorations objectives*

# MRO & Landing Sites: Jezero Crater

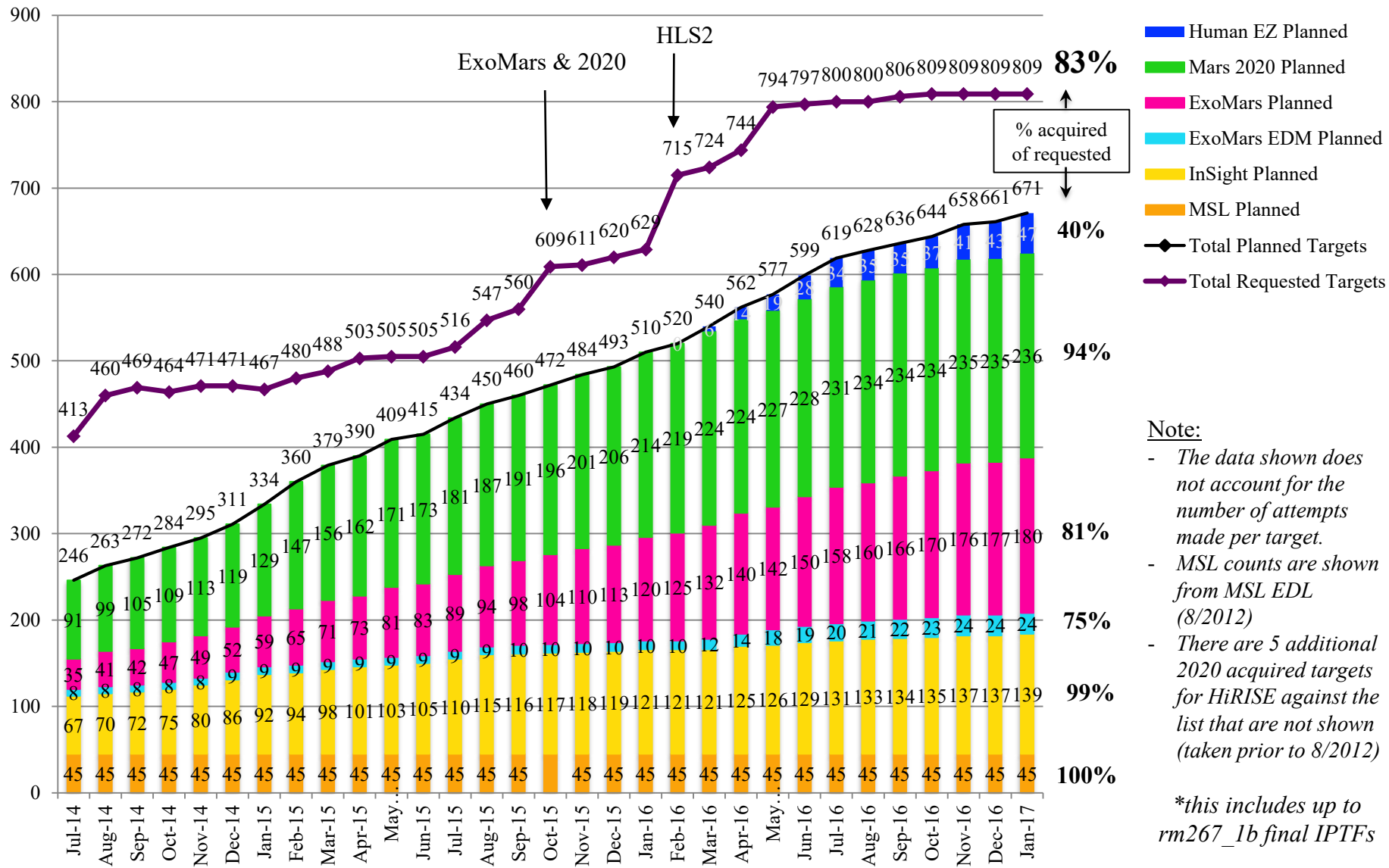


*CRISM / JHUAPL / JPL / NASA*



# Progress on Landing Sites

Number of Targets



# MRO: More to Come!

***MRO continues to operate nominally in its dual-purpose mission of scientific observation and programmatic support***

- ***MRO is working to ensure spacecraft operability through 2023 (2020 Rover Prime Mission)***
  - All-Stellar Mode being developed to preserve IMU lifetime
  - Eclipse power management being instituted to extend battery life
  - Onboard fuel adequate for nominal operations and critical event support (2018 *InSight*, 2020 Rovers)
  - Landing site reconnaissance for *InSight*, 2020 Mars & ExoMars rovers continues
  - Preparing to support EDL and surface relay for *InSight* and Red Dragon – Continuing relay support for MER and MSL
- ***Exciting EM4 mission is in progress with all instruments operating***
  - 1 of 3 CRISM coolers continuing to yield good data in bimonthly cold cycles
  - HiRISE detector aging mitigated by warm-ups
  - MCS, MARCI, CTX, SHARAD show no signs of aging



***Earth (and Moon) as seen by MRO on Nov 20<sup>th</sup> from Mars orbit.***

*Acquired at a range of 205 million km (~200 km/pixel)*







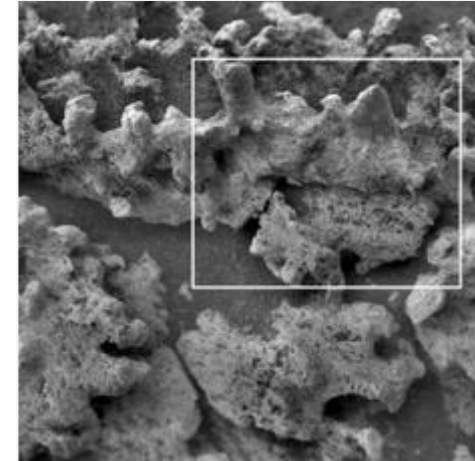
## JEZERO

- Deltaic/lacustrine deposition with Hesperian lava flow and hydrous alteration
- Evidence for hydrous minerals from CRISM, *including carbonates*



## NE SYRTIS

- Extremely ancient igneous, hydrothermal, and sedimentary environments
- High mineralogic diversity with phyllosilicates, sulfates, carbonates, olivine
- Serpentinization and subsurface habitability?

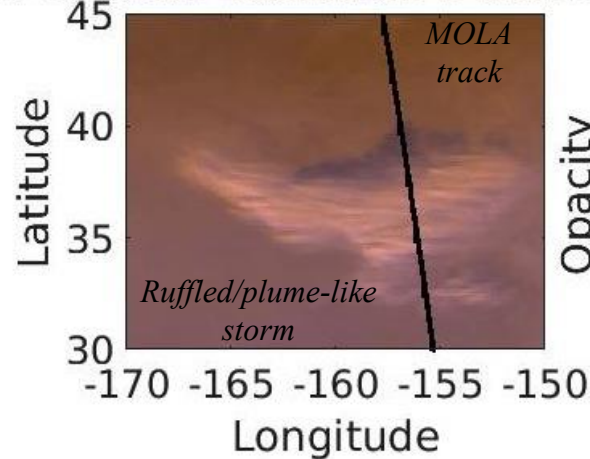


## COLUMBIA HILLS

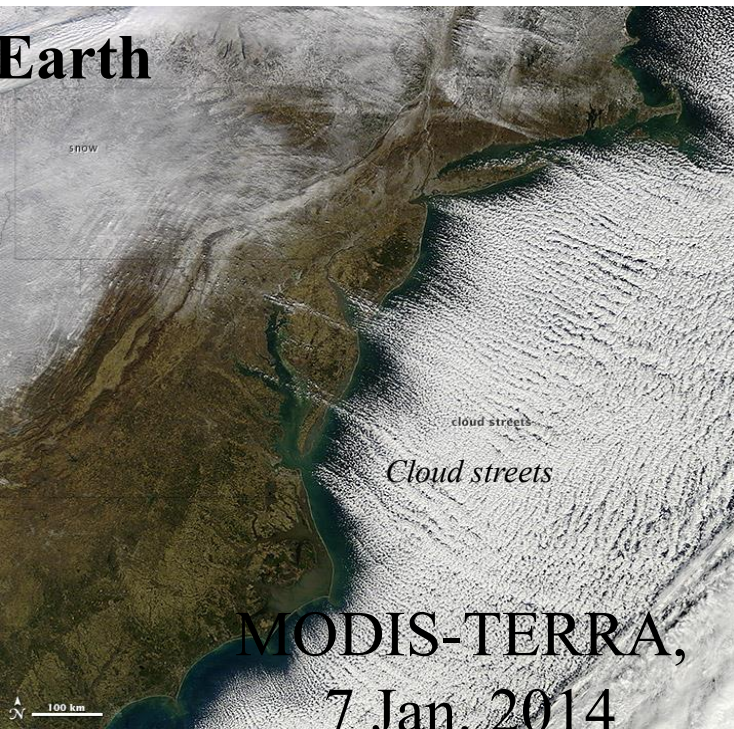
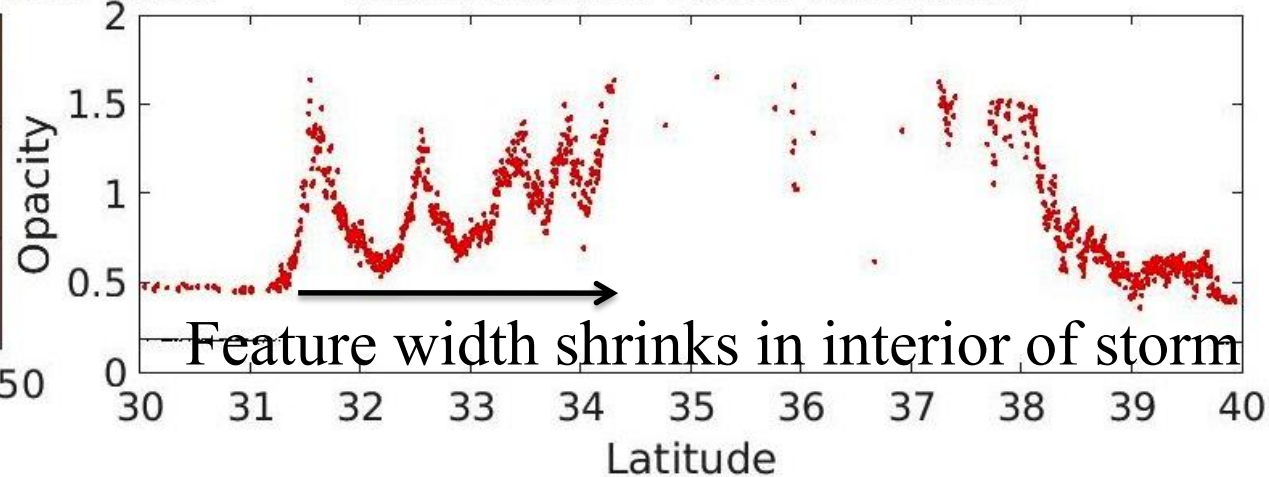
- Carbonate, sulfate, and silica-rich outcrops of possible hydrothermal origin and Hesperian lava flow
- Potential biosignatures identified
- Previously explored by MER

# Structure of One Major Type of Martian Dust Storm Suggests They Are Driven by Strong Winds Behind Cold Fronts

**Mars Orbiter Camera: 3 March 2000**



**Mars Orbiter Laser Altimeter**



Dust storms on Mars have three major types of cloud-top structures in visible imagery. One type (ruffled/plume-like) is commonly observed in a very smooth area with a dusty surface in Mars's northern plains. Using observations from multiple types of instruments on MGS and MRO, it was found that the elongated linear features in these storms strongly resemble cloud streets in the Earth's atmosphere, particularly the type known as "wide, mixed layer rolls." On Earth, these form over bodies of water after the passage of cold fronts. The storms on Mars likely have a similar dynamical origin, with the dusty Martian surface acting like a body of water.

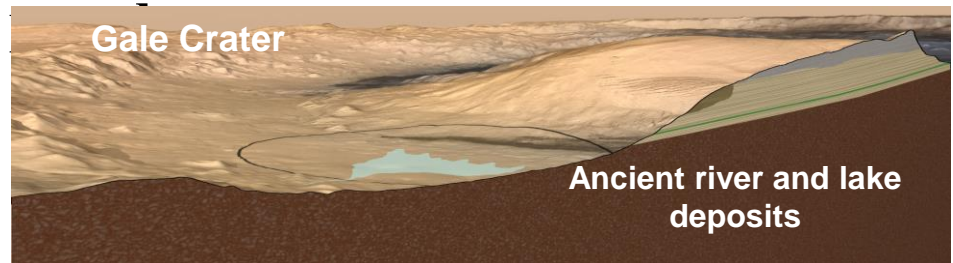
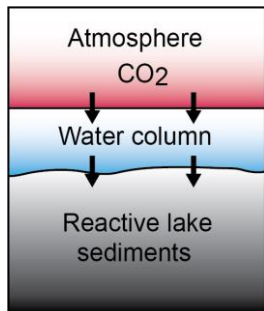
**For more details, see:** N.G. Heavens (2017), Textured Dust Storm Activity in NE Amazonis–SW Arcadia, Mars: Phenomenology and Dynamical Interpretation, accepted by *J. Atmos. Sci.*



# Curiosity: Little CO<sub>2</sub> in Ancient Mars Atmosphere

## Examining minerals in mudstones formed at the bottom of a lake three and half billion years ago on Mars deepens the ‘faint young Sun’

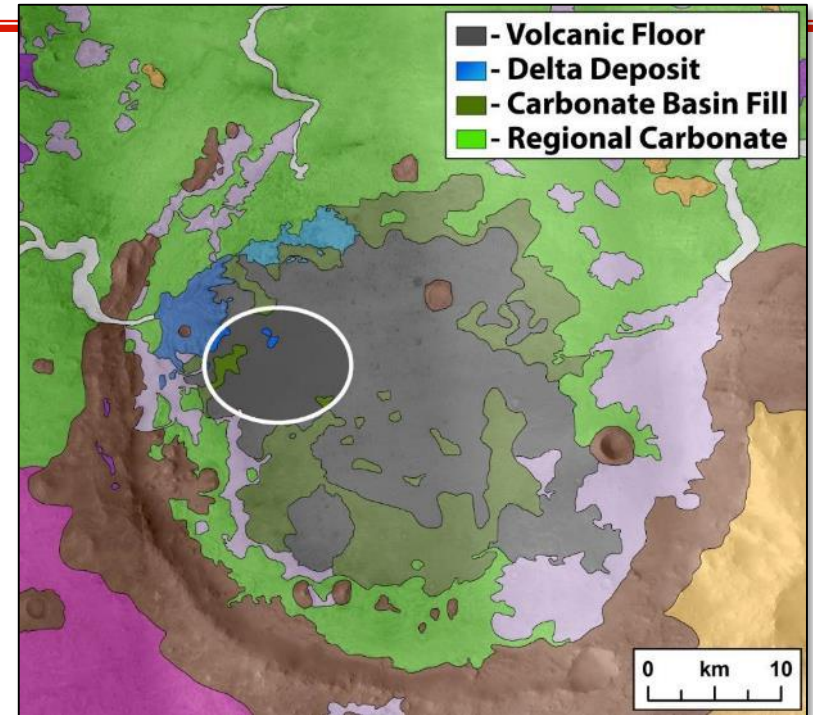
- Satellites have spent decades looking for carbonate minerals on Mars – the remnants of a thick carbon dioxide rich atmosphere thought to have helped keep the planet warmer early in its history than today, despite a dimmer Sun.
- Carbonate minerals are not as common as expected, but may be buried deep within the crust or covered in dust, limiting visibility from space.



- The Mars Science Laboratory (MSL) *Curiosity* Rover sampled ancient lake sediments in Gale Crater containing basaltic minerals that should have reacted to form carbonates if a thick carbon dioxide atmosphere were present. CheMin, an X-ray diffraction instrument on MSL that detects minerals, has not found any carbonates. This limits carbon dioxide to tens to hundreds of times lower than levels required by some Mars climate models in which lakes and rivers can form and flow on the surface without freezing.
- Other ways to keep ancient Mars warm enough to explain evidence of river networks and lakes widespread across the planet in older terrains are needed.

Bristow et al. 2017- *PNAS*

# Jezero Crater

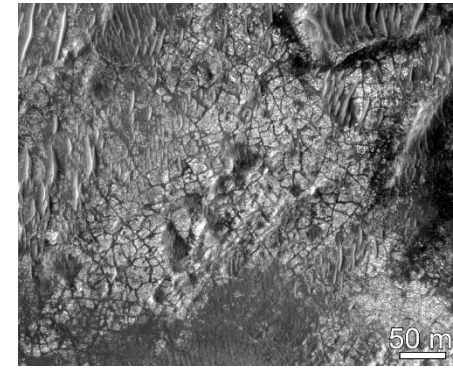
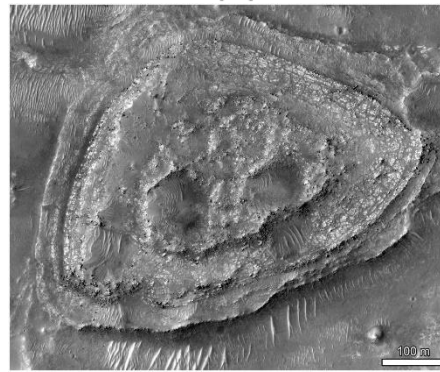
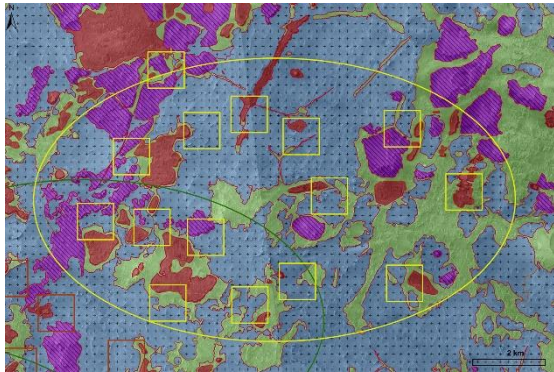
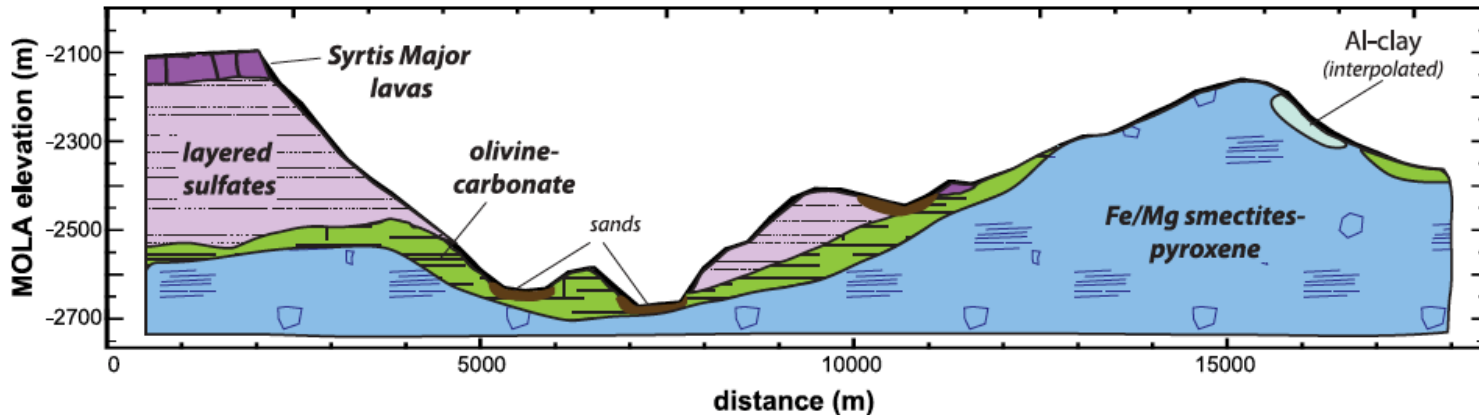


from T. Goudge presentation at LSW3

Positives: Oldest crater lake site (Noachian), with well-defined fine-grained deltaic facies attractive for biosignature investigation. Large, geologically diverse headwaters region. Carbonate bearing unit that may preserve record of ancient climate. Deep open-basin lake.



# NE Syrtis

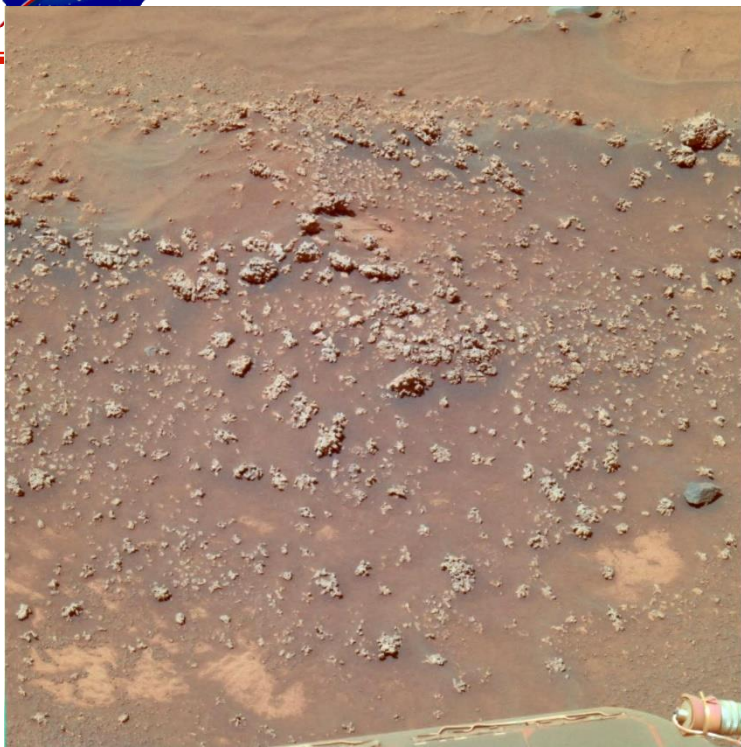


from J. Mustard and M. Bramble at LSW

**Positives:** Lithologic diversity spanning long period of early Mars. Clear, readily accessible stratigraphy within the ellipse. Isidis(?) megabreccia, phyllosilicates, abundant carbonates in stratigraphy (subsurface aquifer?, serpentinizing system?).

**Negatives:** In-ellipse mafic unit may not be volcanic. In-place Syrtis Major lavas probably too distant from ellipse to reach.

# Columbia Hills



from presentations by S. Ruff and R. Arvidson at LSW3

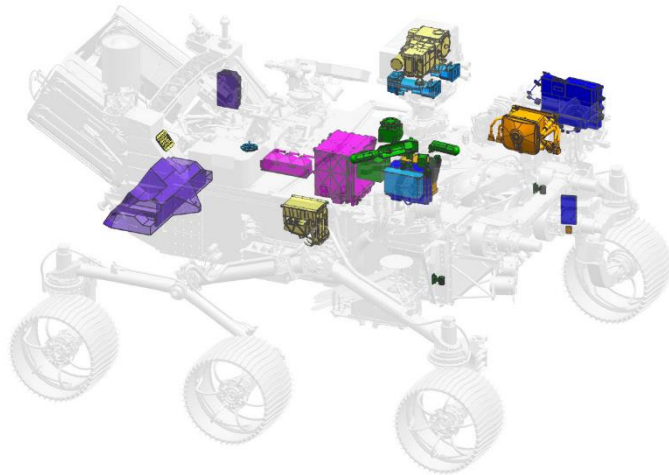
**Positives:** Digitate silica structures possibly from ancient hot spring may be potentially habitable environment with high preservation potential. Diverse igneous suite. Highly rated by RSSB.

**Negatives:** Origin/age of key silica deposit is disputed. Environment is dusty, impeding instrument performance. Site has already been investigated with a capable rover, making it less clear what original contributions Mars 2020 can make in-situ.



## Closing on the Final Landing Site List

1. Project Engineering and Science will continue to work closely together to assess any operational challenges at the highly rated sites.
2. The Mars 2020 PSG and science team will critically assess the pros, cons, uncertainties and risks of the highly rated sites. We will work to develop a science plan at each site (including evaluation of specific regions of interest), and what could be accomplished both in-situ and in terms of cached samples.
3. For Columbia Hills, further assessment of the silica target's origin is being undertaken, as will evaluation of any challenges for its study/collection by the rover. The Mars 2020 PSG will consider how and to what extent the rover's instrument suite can advance scientific knowledge at this previously visited site.



Instrument Key
<b>Mastcam-Z</b> Stereo Imager
<b>MEDA</b> Mars Environmental Measurement
<b>MOXIE</b> In-Situ Oxygen Production
<b>PIXL</b> Microfocus X-ray fluorescence spectrometer
<b>RIMFAX</b> Ground Penetrating Radar
<b>SHERLOC</b> Fluorescence and Raman spectrometer and Visible context imaging
<b>SuperCam</b> LIBS and Raman

Full complement of instruments is still under development, with no reduction in capabilities

Five of seven instruments have had their Critical Design Review (only MEDA and PIXL remain)

Some Updates:

- 1) MEDA schedule is tight
- 2) Good progress is being made resolving SHERLOC (laser) and PIXL (x-ray tube) technical challenges



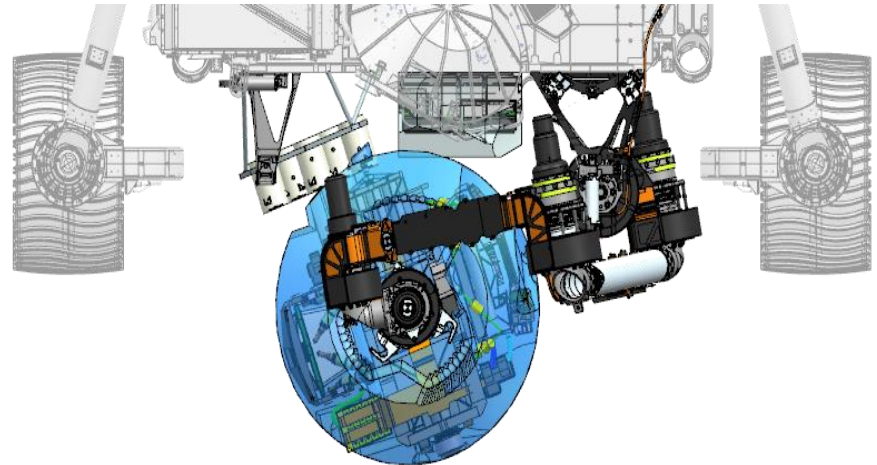
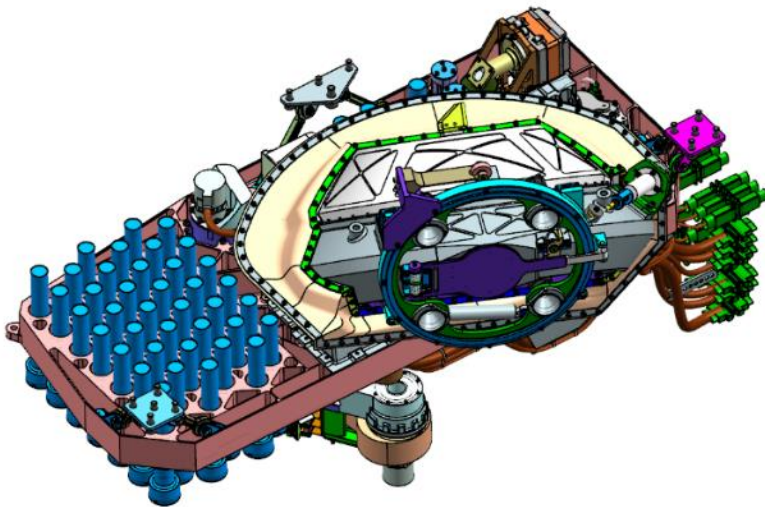
# Sampling and Caching System (SCS)

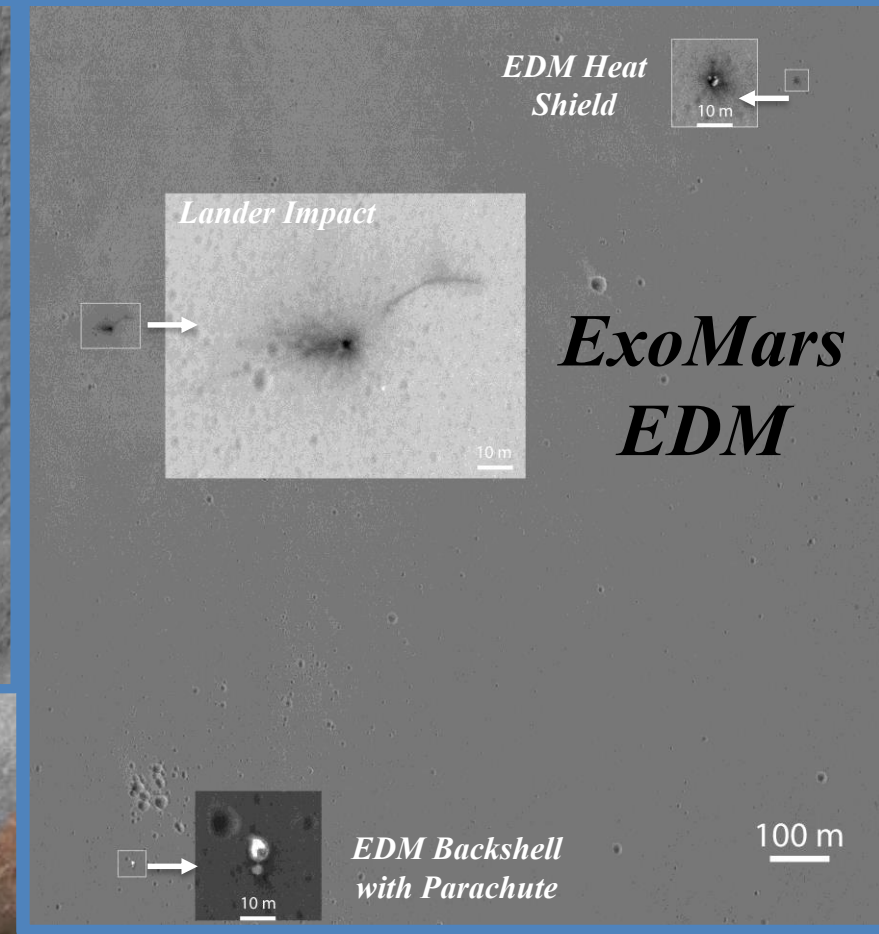
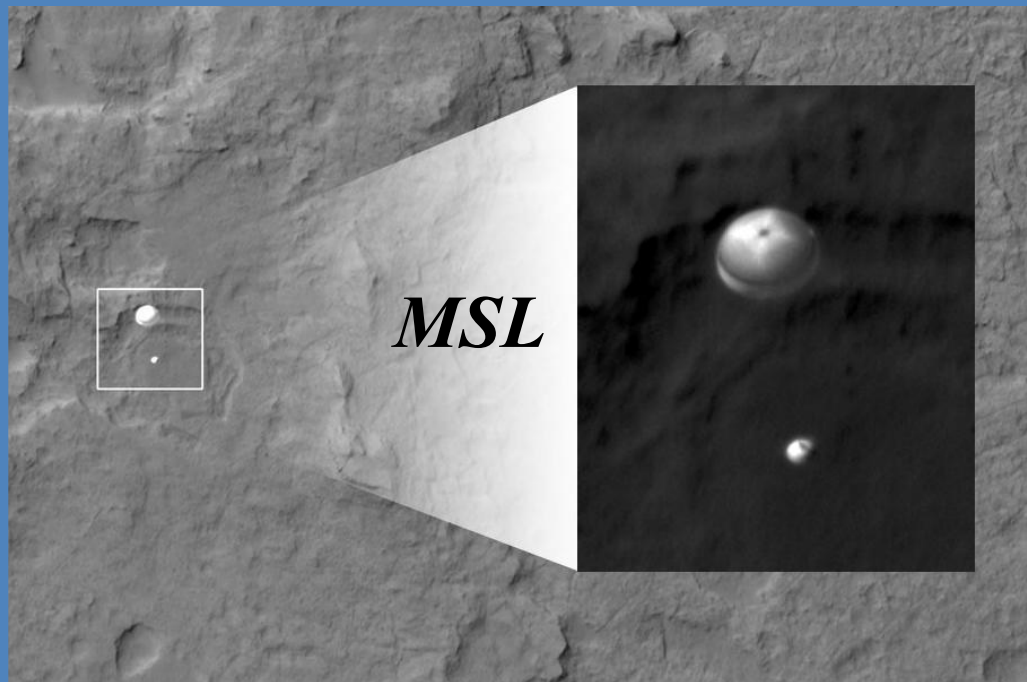
The very complex robotic SCS passed its subsystem CDR earlier this year

Carries ~40 sample tubes (31 + science/engineering spares)

Carries 6 witness tubes (in the 40) from which to acquire round-trip contamination knowledge  
- witness tube design is currently under development

Contamination requirements remain unchanged and good progress is being made in achieving them (e.g., 10 ppb TOC baseline, 40 ppb threshold)





*HiRISE / U. Arizona / JPL / NASA*





## Other Ongoing and Future missions

- Indian space agency - MOM
- Red Dragon ~ 2020
- Chinese program

11 Years in Orbit

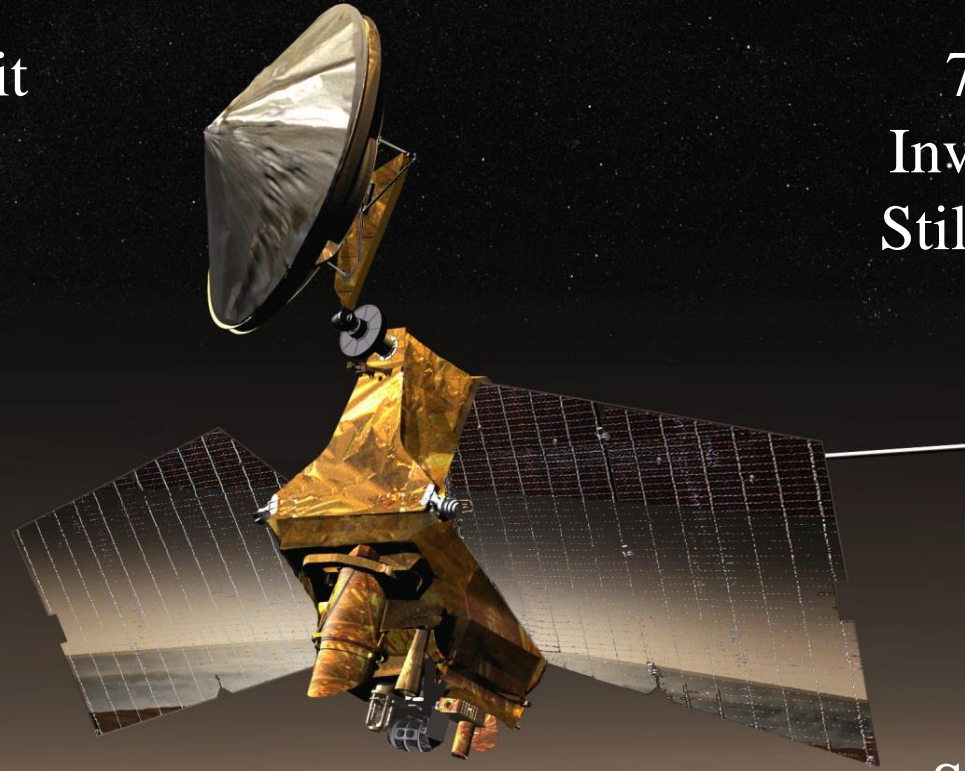
~50,000 orbits

300 Tb of  
Science Data  
Returned

~200 kg of  
Usuable Fuel  
still in the Tank

7 Science  
Investigations  
Still Returning  
Data

More  
Discoveries  
Sure to Come!



## **Mars Reconnaissance Orbiter**

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*Government sponsorship acknowledged.*



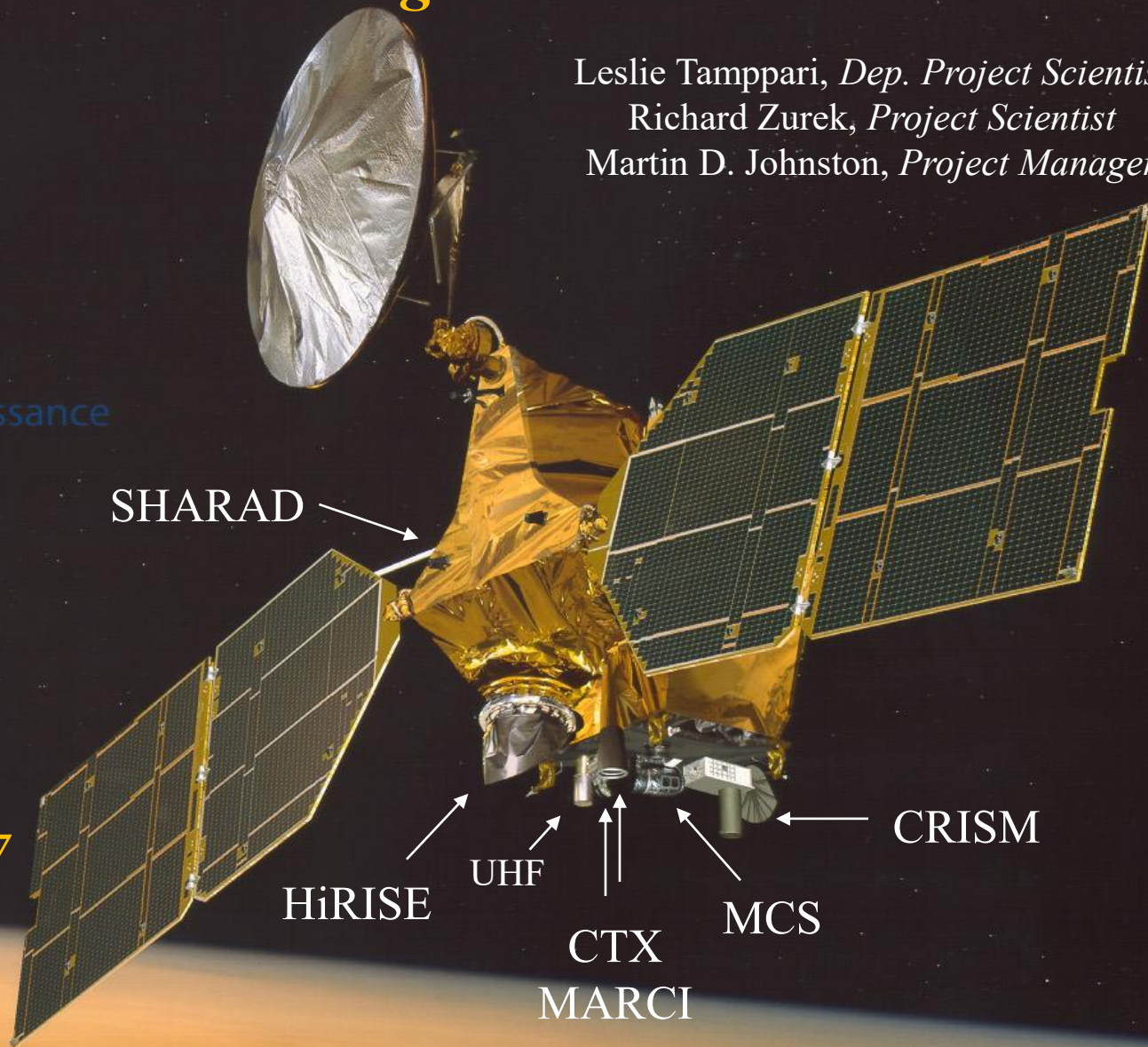
# Backup



# MRO: Continuing Discoveries at Mars

Leslie Tamppari, *Dep. Project Scientist*  
Richard Zurek, *Project Scientist*  
Martin D. Johnston, *Project Manager*

Mars  
Reconnaissance  
Orbiter



Presentation  
to MEPAG

February 23, 2017

CL#17-0802



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